

# Polarized Nucleons: What is the nucleon's spin made off ?

Johannes Blümlein

DESY



1. Introduction
2. QCD Analysis Formalism
3. World Data
4. Parton Distributions with Errors
5.  $\Lambda_{\text{QCD}}$  and  $\alpha_s$
6. Factorization Scheme Invariant Evolution
7. Moments: Comparison with Lattice QCD
8. Integral Relations: Twist 2 & 3
9. Angular Momentum
10. Conclusions

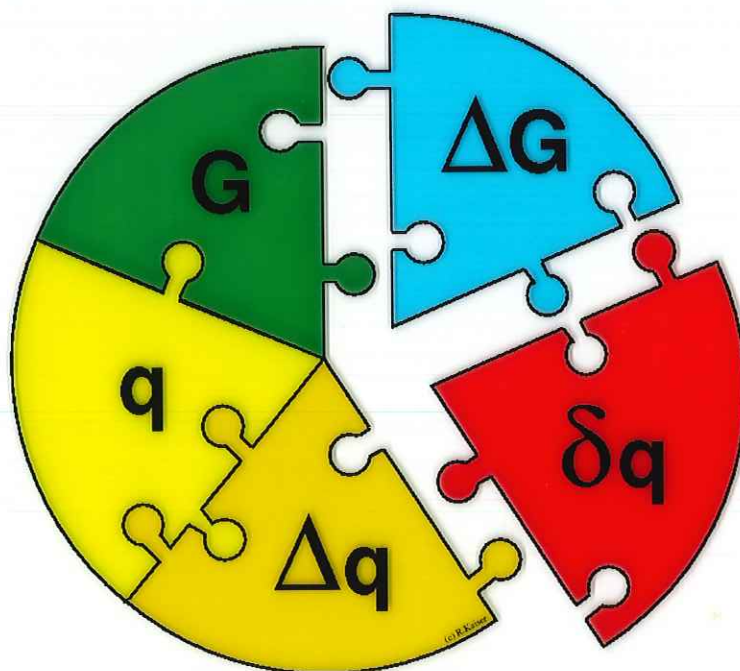
# Introduction

---

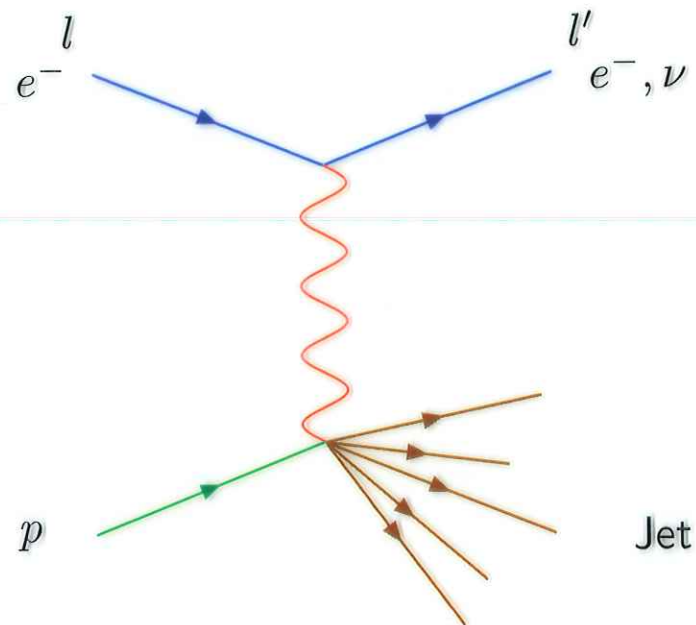
## WHY AND FOR WHICH PURPOSE DO WE STUDY POLARIZED DEEP INELASTIC SCATTERING ?

- Study of short distance structure of nucleon spin
- Test of perturbative QCD in spin sector:  $\Lambda_{QCD}$
- Test of fundamental and less fundamental sum rules
- Does QCD describe polarized nucleons non-perturbatively? (QCD Moments vs. Lattice Moments)

SOLVE THE SPIN PUZZLE!



## DEEPLY INELASTIC SCATTERING



space-like process :

$$q^2 = (l - l')^2 = -Q^2 < 0$$
$$W^2 = (p + q)^2 \geq M_p^2$$

$$x = \frac{Q^2}{2p \cdot q}, \quad y = \frac{p \cdot q}{p \cdot l}$$

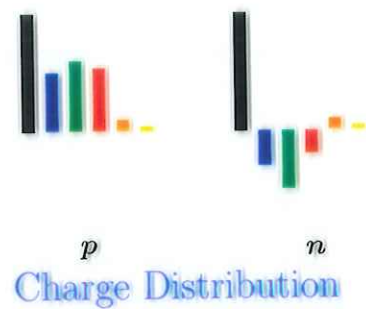
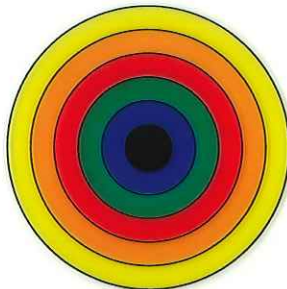
$$0 \leq x, y \leq 1$$

# THE RESOLUTION OF THE NUCLEON MICROSCOPE

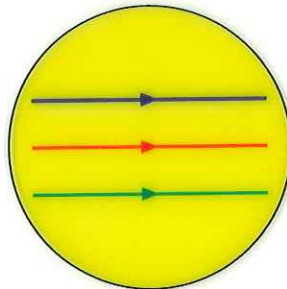
$$\Delta x \sim \frac{1}{|Q|} = \frac{1}{\sqrt{-q^2}}$$

Examples :

$$Q^2 \sim 0.5 \cdot M_p^2$$

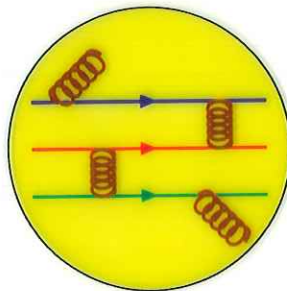


$$Q^2 \sim 3 \cdot M_p^2$$



Scaling

$$Q^2 \sim 10 \dots 500 \cdot M_p^2$$



Violation of Scaling

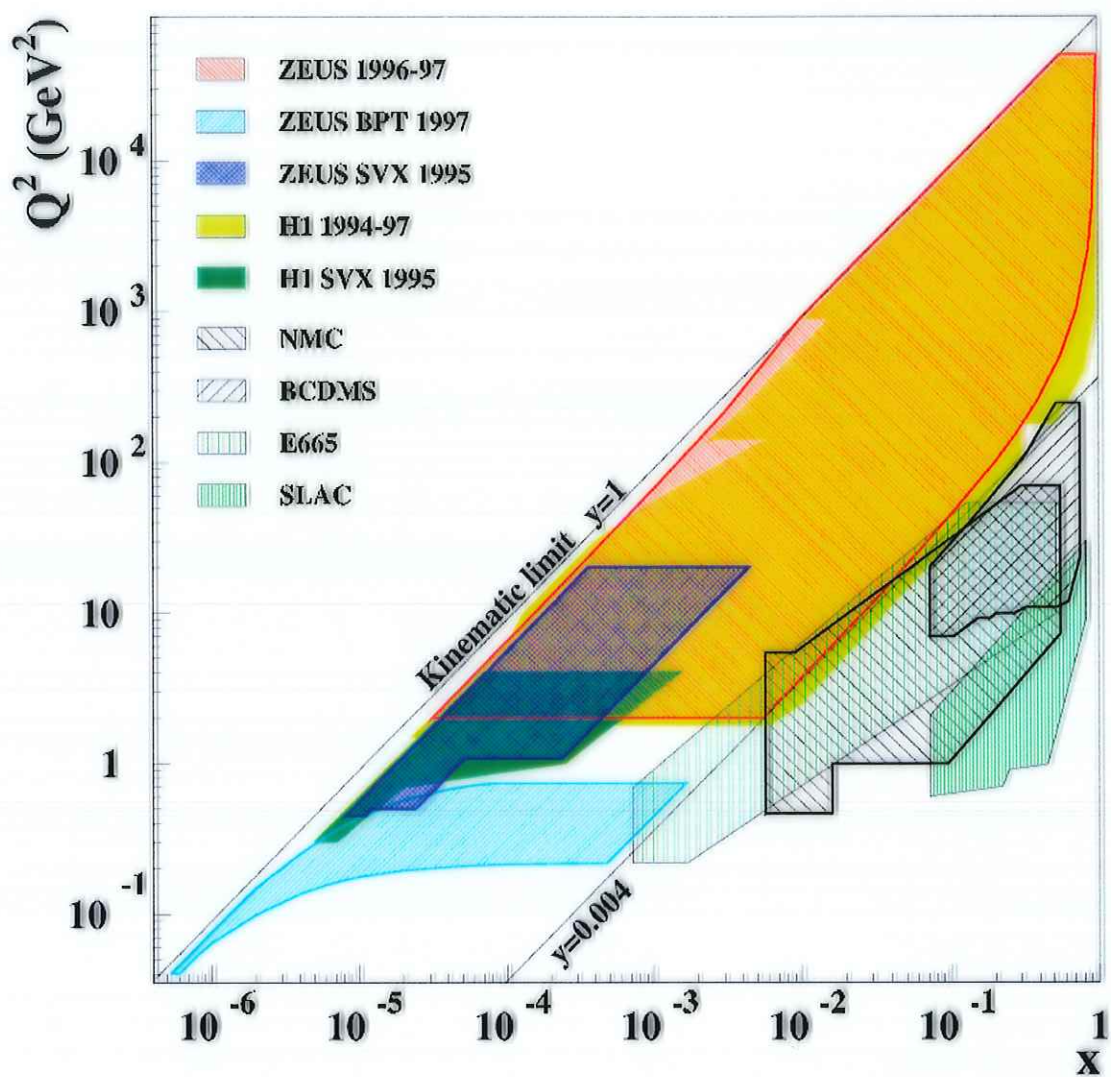
IF THERE ARE NEW COMPOSITENESS SCALES, ONE MAY FIND THEM IN THE FUTURE.

$$Q^2 > 10^4 \text{ GeV}^2,$$

$$1 \text{ GeV}^2 \sim M_p^2$$



## Kinematic Domain



## WHEN IS A PARTON ?

S. DRELL: Infinite Momentum Frame:  $P$  - large

$$\tau_{\text{int}} \ll \tau_{\text{life}}$$

$$\tau_{\text{int}} \sim \frac{1}{q_0} = \frac{4Px}{Q^2(1-x)}$$

$$\tau_{\text{life}} \sim \frac{1}{\sum_i E_i - E} = \frac{2P}{\sum_i (k_{\perp i}^2 + M_i^2)/x_i - M^2} \simeq \frac{2Px(1-x)}{k_{\perp}^2}$$

$$\frac{\tau_{\text{int}}}{\tau_{\text{life}}} = \frac{2k_{\perp}^2}{Q^2(1-x)^2}$$

Stay away from  $x \rightarrow 0$ , since  $xP$  becomes too small.

Stay away from  $x \rightarrow 1$ .

$$Q^2 \gg k_{\perp}^2.$$

## 7. Polarized Nucleons

HOW IS THE NUCLEON SPIN DISTRIBUTED OVER THE PARTONS?

$$S_n = \frac{1}{2} [\Delta(u + \bar{u}) + \Delta(d + \bar{d}) + \Delta(s + \bar{s})] + \Delta G + L_q + L_g$$

$$S_n = \frac{1}{2}$$

$$\Delta\Sigma = 0.138 \pm 0.082, \quad (0.150 \pm 0.061)$$

$$\Delta G = 1.026 \pm 0.554, \quad (0.931 \pm 0.679)$$

EMC, 1987: THE NUCLEON SPIN IS NOT THE SUM OF THE LIGHT QUARK SPINS.

MEASURE:

POLARIZED PARTON DENSITIES:  $\Delta q_i, \Delta G$

HOW CAN ONE ACCESS THE PARTON ANGULAR MOMENTUM ?

POLARIZED HEAVY FLAVOR CONTRIBUTIONS.

• POLARIZED STRUCTURE FUNCTIONS CONTAIN ALSO TWIST 3 CONTRIBUTIONS.

HOW TO UNFOLD THESE TERMS ?



## Motivation

---

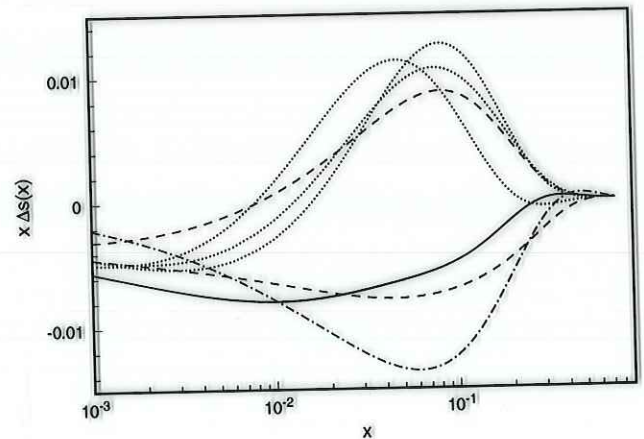
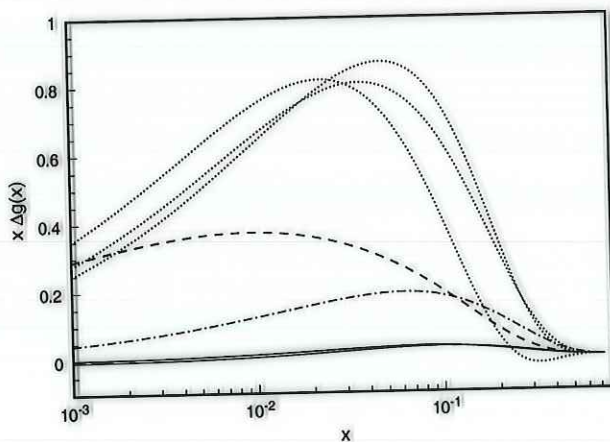
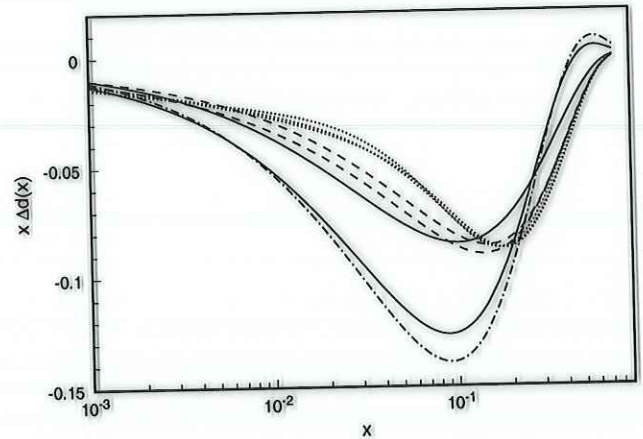
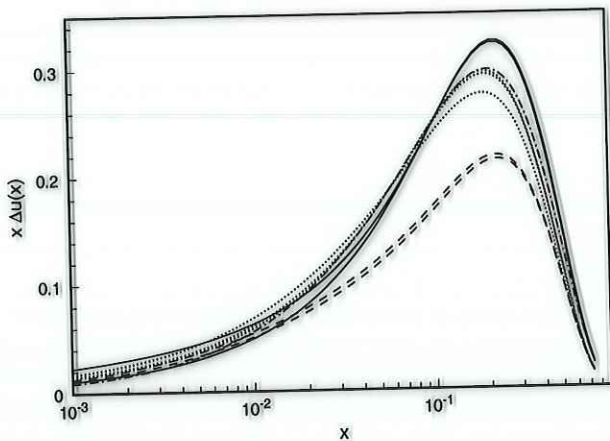
- A number of QCD analyses for polarized data performed so far :
  - T.Gehrmann and W.J.Stirling (GS), Phys.Rev.**D53**(1996)6100.
  - G.Altarelli et al. (ABFR), Nucl.Phys.**B496**(1997)337.
  - Y.Goto et al. (AAC), Phys.Rev.**D62**(2000)034017; London  $\nu$ -Factory Conference to appear. **HIRAI, KUMANO, SAITO et al.**
  - M.Glück et al. (GRSV), Phys.Rev.**D63**(2001)094005.
  - E.Leader et al. (LSS), Eur.Phys.J.**C23**(2002)479.
  - J. Blümlein and H. Böttcher, Nucl. Phys. **B636** (2002) 225
  
  - E154 Collaboration, Phys.Lett.**B405**(1997)180.
  - SMC Collaboration, Phys.Rev.**D58**(1998)112002.

However, **no reliable parametrization of the error bands** for the polarized parton densities are given.  
(most of the cases.)

- We aim at parametrizations of polarized densities and their fully correlated  $1\sigma$  error bands which are directly applicable to determine 'experimental' errors for other polarized observables.
- Such an analysis has a value of its own within the framework of spin physics in order to understand the spin puzzle.
- Comparison of QCD analysis results with results from recent lattice simulations concerning both QCD parameters and low order moments.

# Comparison Polarized Parton Densities

- Compilation by G. A. Ladinsky: (at  $Q = 15 \text{ GeV}$ )  
**Ref.:** Proc. of the Workshop on Prospects of SPIN PHYSICS  
 at HERA p.285, DESY 95-200, MSU-51120, hep-ph/9601287.



⇒ solid lines: N-94 sets 1 and 2

**Ref.:** Nadolsky, Z.Phys. C63 (1994) 601.

⇒ dashed-dotted line: BBS-95

**Ref.:** Brodsky, Burkardt and Schmidt, Nucl.Phys. B441 (1995) 197.

⇒ dotted lines: GS-95 sets A, B, and C

**Ref.:** Gehrmann and Stirling, Z.Phys. C65 (1995) 461.

⇒ dashed lines: GRV-95 standard and valence scenarios

**Ref.:** Glück, Reya, and Vogelsang, Phys.Lett. B359 (1995) 201.



## Evolution in MELLIN space

---

- The polarized structure function  $g_1(x, Q^2)$  represented in terms of a MELLIN convolution of polarized parton densities  $\Delta f_j$  and Wilson coefficients  $\Delta C_j$ :

$$\begin{aligned}
 g_1(x, Q^2) = & \frac{1}{2} \sum_{j=1}^{N_f} e_j^2 \int_x^1 \frac{dz}{z} \left[ \frac{1}{N_f} \Delta\Sigma \left( \frac{x}{z}, \mu_f^2 \right) \Delta C_q^S \left( z, \frac{Q^2}{\mu_f^2} \right) \right. \\
 & + \Delta G \left( \frac{x}{z}, \mu_f^2 \right) \Delta C_G \left( z, \frac{Q^2}{\mu_f^2} \right) \\
 & \left. + \Delta q_j^{NS} \left( \frac{x}{z}, \mu_f^2 \right) \Delta C_q^{NS} \left( z, \frac{Q^2}{\mu_f^2} \right) \right],
 \end{aligned}$$

with the **singlet density**  $\Delta\Sigma$

$$\Delta\Sigma \left( z, \mu_f^2 \right) = \sum_{j=1}^{N_f} \left[ \Delta q_j \left( z, \mu_f^2 \right) + \Delta \bar{q}_j \left( z, \mu_f^2 \right) \right],$$

the **gluon density**  $\Delta G$ ,

the **non-singlet density**  $\Delta q_j^{NS}$

$$\begin{aligned}
 \Delta q_j^{NS} \left( z, \mu_f^2 \right) = & \Delta q_j \left( z, \mu_f^2 \right) + \Delta \bar{q}_j \left( z, \mu_f^2 \right) \\
 & - \frac{1}{N_f} \Delta\Sigma \left( z, \mu_f^2 \right),
 \end{aligned}$$

and the factorization scale  $\mu_f$ .

- The above quantities also depend on the **renormalization scale**  $\mu_r$  of the strong coupling constant  $a_s(\mu_r^2) = g_s^2(\mu_r^2)/(16\pi^2)$ . The observable  $g_1(x, Q^2)$  is independent of the choice of both scales.

## Evolution in MELLIN space (cont'd)

- The evolution equations are given by

$$\frac{\partial \Delta q_i^{\text{NS}}(x, Q^2)}{\partial \log Q^2} = \Delta P_{\text{NS}}^-(x, a_s) \otimes \Delta q_i^{\text{NS}}(x, Q^2)$$

$$\frac{\partial}{\partial \log Q^2} \begin{pmatrix} \Delta \Sigma(x, Q^2) \\ \Delta G(x, Q^2) \end{pmatrix} = \Delta \mathbf{P}(x, a_s) \otimes \begin{pmatrix} \Delta \Sigma(x, Q^2) \\ \Delta G(x, Q^2) \end{pmatrix}$$

with

$$\Delta P_{\text{NS}}^-(x, a_s) = a_s \Delta P_{\text{NS}}^{(0)}(x) + a_s^2 \Delta P_{\text{NS}}^{-(1)}(x) + \mathcal{O}(a_s^3)$$

$$\Delta \mathbf{P}(x, a_s) \equiv \begin{pmatrix} \Delta P_{qq}(x, Q^2) & \Delta P_{qg}(x, Q^2) \\ \Delta P_{gq}(x, Q^2) & \Delta P_{gg}(x, Q^2) \end{pmatrix}$$

$$= a_s \Delta \mathbf{P}^{(0)}(x) + a_s^2 \Delta \mathbf{P}^{(1)}(x) + \mathcal{O}(a_s^3)$$

and  $\otimes$  the MELLIN convolution

$$[A \otimes B](x) = \int_0^1 dx_1 dx_2 \delta(x - x_1 x_2) A(x_1) B(x_2)$$

- The polarized Wilson coefficient functions  $\Delta C_i(x, \alpha_s(Q^2))$  and the polarized splitting functions  $\Delta P_{ij}(x, \alpha_s(Q^2))$  are known in the  $\overline{\text{MS}}$  scheme up to NLO. [W.L. van Neerven and E.B. Zijlstra, Nucl. Phys. B417 (1994) 61, R. Mertig and W.L. van Neerven, Z. Phys. C70 (1996) 637, W. Vogelsang, Phys. Rev. D54 (1996) 2023]



A complete NLO QCD Analysis possible.

## Evolution in MELLIN space (cont'd)

---

- $a_s(\mu_r)$  is obtained as the solution of

$$\mu_r^2 \frac{da_s(\mu_r^2)}{d\mu_r^2} = -\beta_0 a_s^2(\mu_r^2) - \beta_1 a_s^3(\mu_r^2) + \mathcal{O}(a_s^4),$$

where the coefficients of the  $\beta$ -function are given by (in the  $\overline{\text{MS}}$  scheme)

$$\beta_0 = \frac{11}{3}C_A - \frac{4}{3}T_F N_f,$$

$$\beta_1 = \frac{34}{3}C_A^2 - \frac{20}{3}C_A T_F N_f - 4C_F T_F N_f,$$

and

$$C_A = 3, \quad T_F = 1/2, \quad C_F = 4/3.$$

- $\Lambda_{\text{QCD}}^{\overline{\text{MS}}}$  is given by:

$$\Lambda_{\text{QCD}}^{\overline{\text{MS}}} = \mu_r \exp \left\{ -\frac{1}{2} \left[ \frac{1}{\beta_0 a_s(\mu_r^2)} - \frac{\beta_1}{\beta_0^2} \log \left( \frac{1}{\beta_0 a_s(\mu_r^2)} + \frac{\beta_1}{\beta_0} \right) \right] \right\}.$$

⇒ We extract  $\Lambda_{\text{QCD}}^{(4)}$  from the data and choose  $N_f = 4$  whereas the polarized structure function  $g_1(x, Q^2)$  is presented using only the **three light flavors**.



## Evolution in MELLIN space (cont'd)

---

- The evolution equations are solved analytically in **MELLIN- $N$**  space:

→ A MELLIN-transformation is performed

$$\mathbf{M}[f](N) = \int_0^1 dx x^{N-1} f(x), \quad N \in \mathbf{N},$$

which turns the MELLIN convolution  $\otimes$  into an ordinary product.

- The non-singlet solution:

$$\begin{aligned} \Delta q^{\text{NS}}(N, a_s) &= \left( \frac{a_s}{a_0} \right)^{-P_{\text{NS}}^{(0)}/\beta_0} \left[ 1 - \frac{1}{\beta_0} (a_s - a_0) \right. \\ &\quad \left. \times \left( P_{\text{NS}}^{-(1)} - \frac{\beta_1}{\beta_0} P_{\text{NS}}^{(0)} \right) \right] \Delta q^{\text{NS}}(N, a_0) \end{aligned}$$

and the singlet solution:

$$\begin{aligned} \begin{pmatrix} \Delta \Sigma(N, a_s) \\ \Delta G(N, a_s) \end{pmatrix} &= [\mathbf{1} + a_s \mathbf{U}_1(N)] \mathbf{L}(N, a_s, a_0) [\mathbf{1} - a_0 \mathbf{U}_1(N)] \\ &\quad \times \begin{pmatrix} \Delta \Sigma(N, a_0) \\ \Delta G(N, a_0) \end{pmatrix}, \end{aligned}$$

where  $a_s = a_s(Q^2)$  and  $a_0 = a_s(Q_0^2)$ .

⇒ The input and the evolution parts factorize.

[W.Furmanski and R.Petronzio, Z.Phys.**C11**(1982)293, M.Glück, E.Reya, and A.Vogt, Z.Phys.**C48**(1990)471, J.Blümlein and A.Vogt, Phys.Rev.**D58**(1998)014020.]

## Evolution in MELLIN space (cont'd)

---

- The **Leading Order** singlet evolution matrix is given by

$$\mathbf{L}(a_s, a_0, N) = \mathbf{e}_-(N) \left( \frac{a_s}{a_0} \right)^{-r_-(N)} + \mathbf{e}_+(N) \left( \frac{a_s}{a_0} \right)^{-r_+(N)}$$

with the eigenvalues

$$r_{\pm} = \frac{1}{\beta_0} \left[ \text{tr}(\mathbf{P}^{(0)}) \pm \sqrt{\text{tr}(\mathbf{P}^{(0)})^2 - \det_2(\mathbf{P}^{(0)})} \right]$$

and the eigenvectors

$$\mathbf{e}_{\pm} = \frac{\mathbf{P}^{(0)}/\beta_0 - r_{\mp} \mathbf{1}}{r_{\pm} - r_{\mp}}.$$

- The **Next-to-Leading Order** singlet solution is obtained from the LO singlet solution through the matrix  $\mathbf{U}_1(N)$

$$\begin{aligned} \mathbf{U}_1(N) = & -\mathbf{e}_- \mathbf{R}_1 \mathbf{e}_- - \mathbf{e}_+ \mathbf{R}_1 \mathbf{e}_+ + \frac{\mathbf{e}_+ \mathbf{R}_1 \mathbf{e}_-}{r_- - r_+ - 1} \\ & + \frac{\mathbf{e}_- \mathbf{R}_1 \mathbf{e}_+}{r_+ - r_- - 1} \end{aligned}$$

with

$$\mathbf{R}_1 = [\mathbf{P}^{(1)} - (\beta_1/\beta_0)\mathbf{P}^{(0)}]/\beta_0.$$



## Evolution in MELLIN space (cont'd)

---

- The input densities

$$\Delta\Sigma(N, a_0), \Delta G(N, a_0), \text{ and } \Delta q_i^{NS}(N, a_0)$$

are evolved to the scale  $Q^2$ , respectively to the coupling  $\alpha_s(Q^2)$ . An **inverse MELLIN–transformation** to  $x$ –space is then performed by a **contour integral** in the complex plane around all singularities:

$$\Delta f(x) = \frac{1}{\pi} \int_0^\infty dz \operatorname{Im} \left[ \exp(i\phi) x^{-c(z)} \Delta f[c(z)] \right].$$

(Path:  $c(z) = c_1 + \rho[\cos(\phi) + i \sin(\phi)]$ , with  $c_1 = 1.1$ ,  $\rho \geq 0$ , and  $\phi = \frac{3}{4}\pi$ .)

- The function  $\Delta f(x)$  finally depends on the parameters of the parton distributions chosen at the input scale  $Q_0^2$  and on  $\Lambda_{QCD}$ . **These parameters are determined by the fit to the data.**

## Parametrization

---

- General choice for the parametrization of the polarized parton distributions at  $Q_0^2$ :

$$x\Delta q_i(x, Q_0^2) = \eta_i A_i x^{a_i} (1-x)^{b_i} (1 + \gamma_i x + \rho_i x^{\frac{1}{2}})$$

- Normalization:

$$A_i^{-1} = \left( 1 + \gamma_i \frac{a_i}{a_i + b_i + 1} \right) \frac{\Gamma(a_i)\Gamma(b_i + 1)}{\Gamma(a_i + b_i + 1)} + \rho_i \frac{\Gamma(a_i + 0.5)\Gamma(b_i + 1)}{\Gamma(a_i + b_i + 1.5)}$$

such that  $\int_0^1 dx \Delta q_i(x, Q_0^2) = \eta_i$

are the first moment of  $\Delta q_i(x, Q_0^2)$ .

- The polarized parton distributions to be fitted are:

$$\Delta u_v, \Delta d_v, \Delta \bar{q}, \Delta G,$$

where the index  $v$  denotes the *valence* quark.

Note :  $\Delta q + \Delta \bar{q} = \Delta q_v + 2\Delta \bar{q}$ .

## Choice of Parameters

---

- Parameters which have been **fixed** since the data do not constrain those parameters well enough:

– For  $\Delta u_v$  and  $\Delta d_v$ :  $\gamma$

$$x\Delta q_i(x, Q_0^2) = \eta_i A_i x^{a_i} (1-x)^{b_i} (1 + \gamma_i x)$$

– For  $\Delta \bar{q}$  and  $\Delta G$ :  $b$

$$x\Delta q_i(x, Q_0^2) = \eta_i A_i x^{a_i} (1-x)^{b_i}$$

- Relations adopted between the parameters  $a_i$  and  $b_i$  for  $\Delta \bar{q}$  and  $\Delta G$ :

$$a_G = a_{\bar{q}} + C, \quad \text{with } 0.5 < C < 1.0.$$

$$\left( \frac{b_{\bar{q}}}{b_G} \right)^{pol} = \left( \frac{b_{\bar{q}}}{b_G} \right)^{unpol}$$

⇒ Essential to respect **Positivity** for  $\Delta \bar{q}$  and  $\Delta G$ .

- No Positivity** constraint assumed for  $\Delta u_v$  and  $\Delta d_v$ .

⇒ Finally 7 parameters are left free to be determined in the fit. In addition  $\Lambda_{QCD}$  is fitted. → (7 + 1)

Note:

$$x\Delta q_i(x, Q_0^2) = \eta_i A_i x^{a_i} (1-x)^{b_i} (1 + \gamma_i x + \rho_i x^{\frac{1}{2}})$$



## The World Data

---

Published Experimental Data above  $Q^2 = 1.0 \text{ GeV}^2$

Experiment	$x$ -range	$Q^2$ -range [ $\text{GeV}^2$ ]	number of data points	
			$g_1/F_1$ or $A_1$	$g_1$
E143(p)	0.027 – 0.749	1.17 – 9.52	82	28
HERMES(p)	0.028 – 0.660	1.13 – 7.46	39	39
E155(p)	0.015 – 0.750	1.22 – 34.72	24	24
SMC(p)	0.005 – 0.480	1.30 – 58.0	59	12
EMC(p)	0.015 – 0.466	3.50 – 29.5	10	10
<i>proton</i>			<b>214</b>	<b>113</b>
E143(d)	0.027 – 0.749	1.17 – 9.52	82	28
E155(d)	0.015 – 0.750	1.22 – 34.79	24	24
SMC(d)	0.005 – 0.479	1.30 – 54.8	65	12
<i>deuteron</i>			<b>171</b>	<b>64</b>
E142(n)	0.035 – 0.466	1.10 – 5.50	30	8
HERMES(n)	0.033 – 0.464	1.22 – 5.25	9	9
E154(n)	0.017 – 0.564	1.20 – 15.0	11	17
<i>neutron</i>			<b>50</b>	<b>34</b>
<i>total</i>			<b>435</b>	<b>211</b>

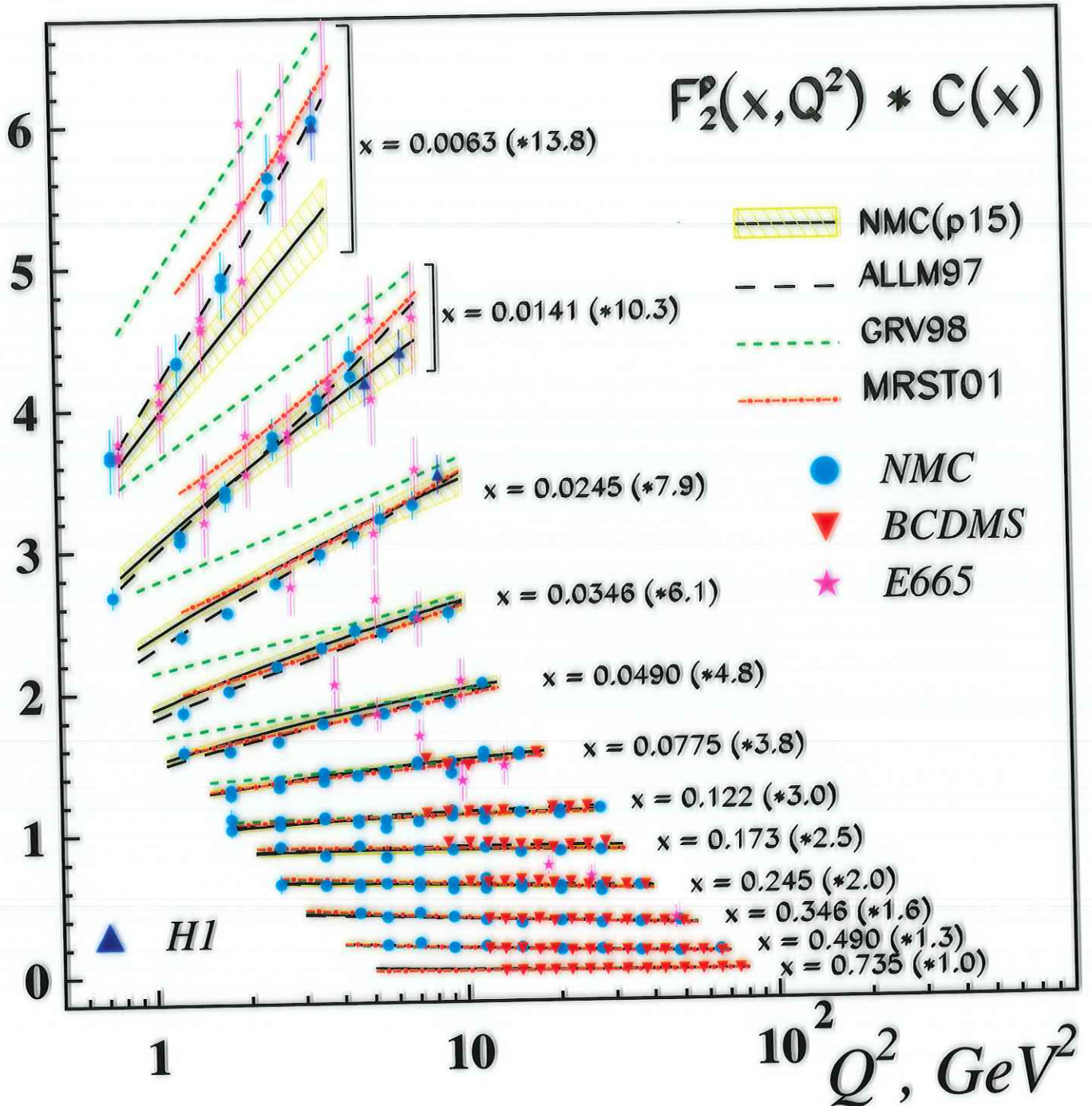
$$g_1/F_1 \approx \frac{1}{(1 + \gamma^2)} A_1, \quad \text{where } \gamma^2 = Q^2/\nu^2$$

$$F_1 = \frac{(1 + \gamma^2)}{2x(1 + R)} F_2$$

$F_2$ -Parametrization: NMC, M. Arneodo et al., Phys. Lett. **B364** (1995) 107.

$R$ -Parametrization: SLAC, L. Withlow et al., Phys. Lett. **B250** (1990) 193.

# $F_2^p$ – Comparison Data / Parameterization

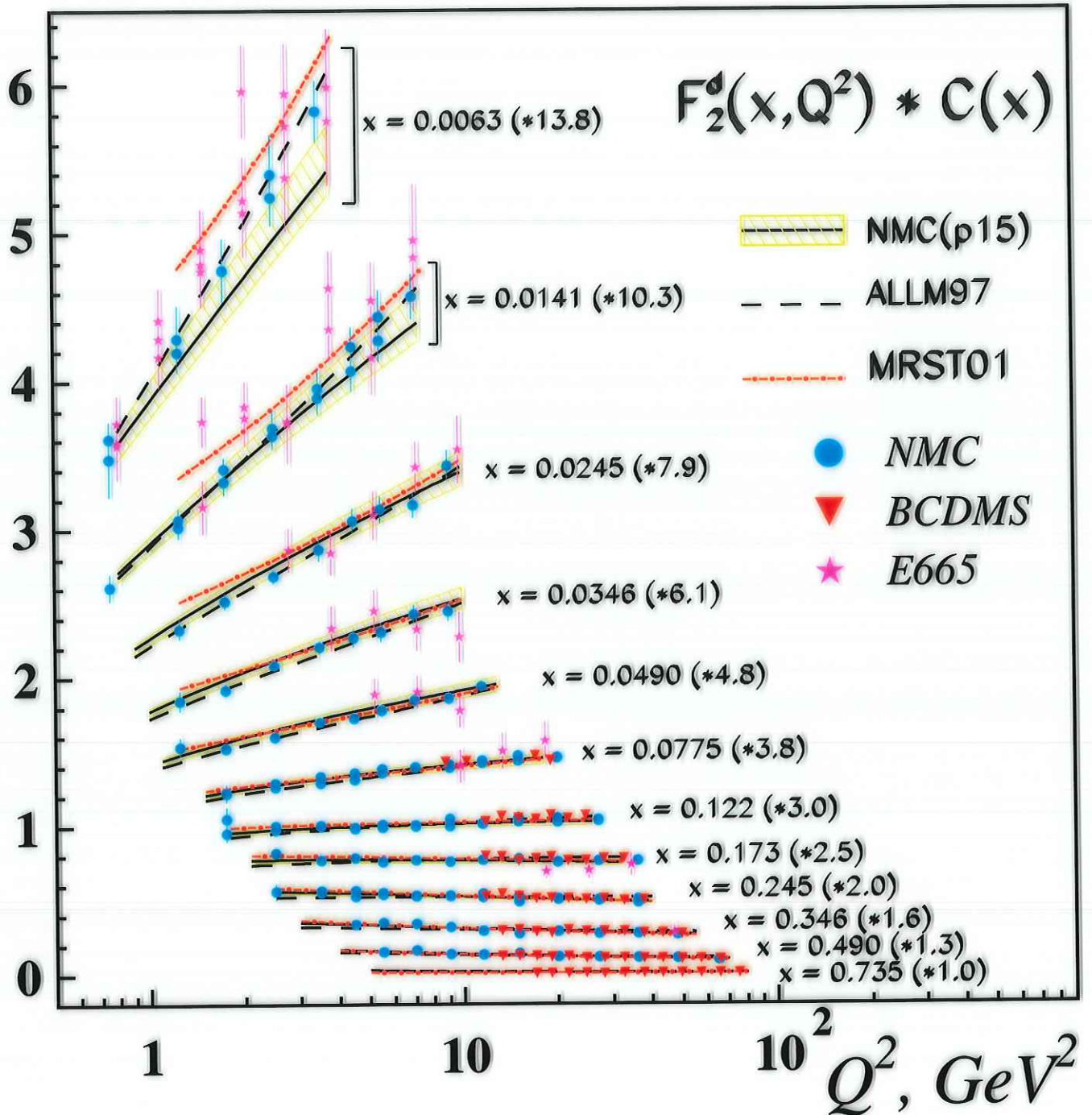


⇒ Question: Which Parameterization to be used?

⇒ Answer: ALLM97



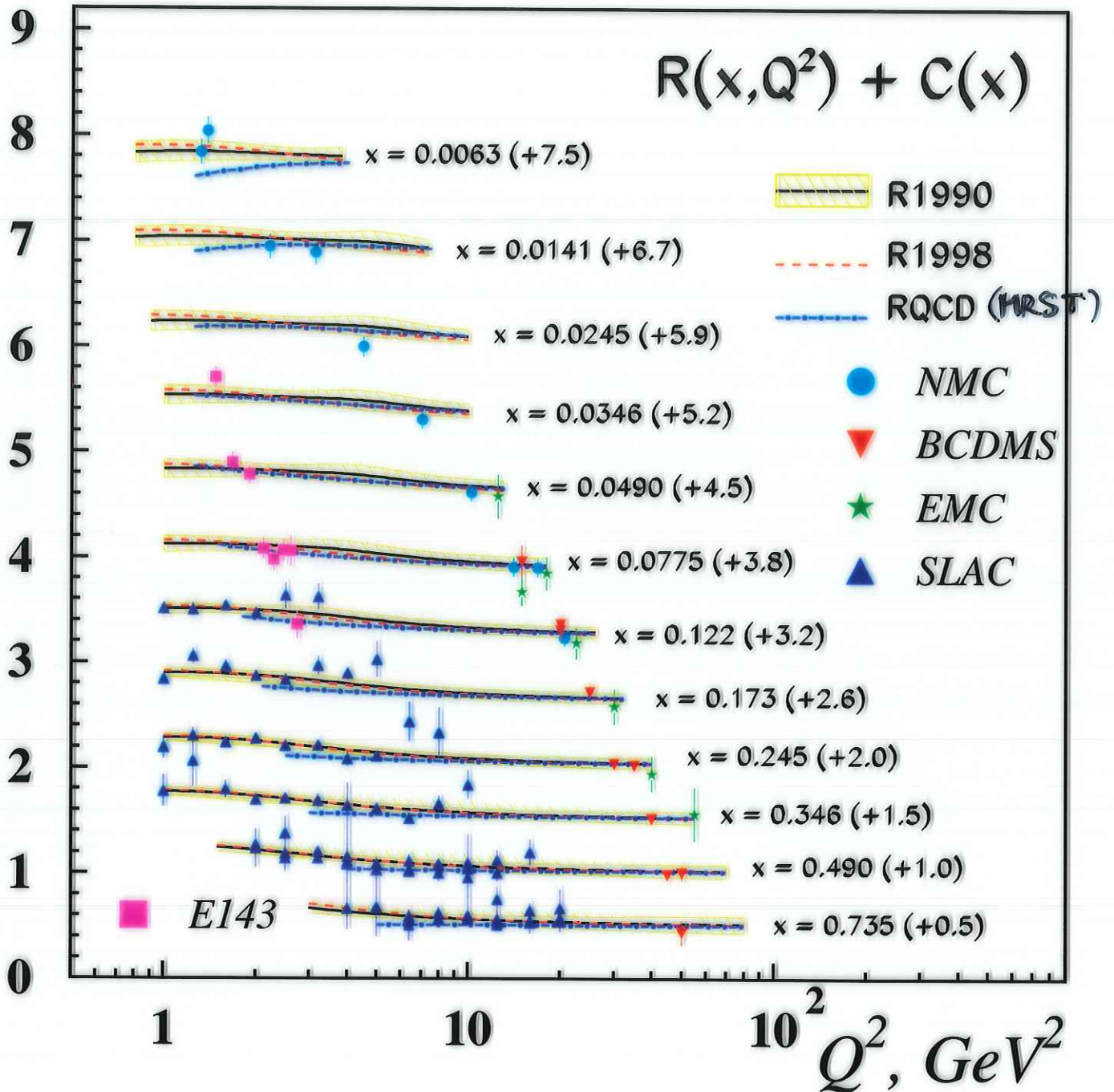
# $F_2^d$ – Comparison Data / Parameterization



⇒ Question: Which Parameterization to be used?

⇒ Answer: ALLM97

# R – Comparison Data / Parameterization

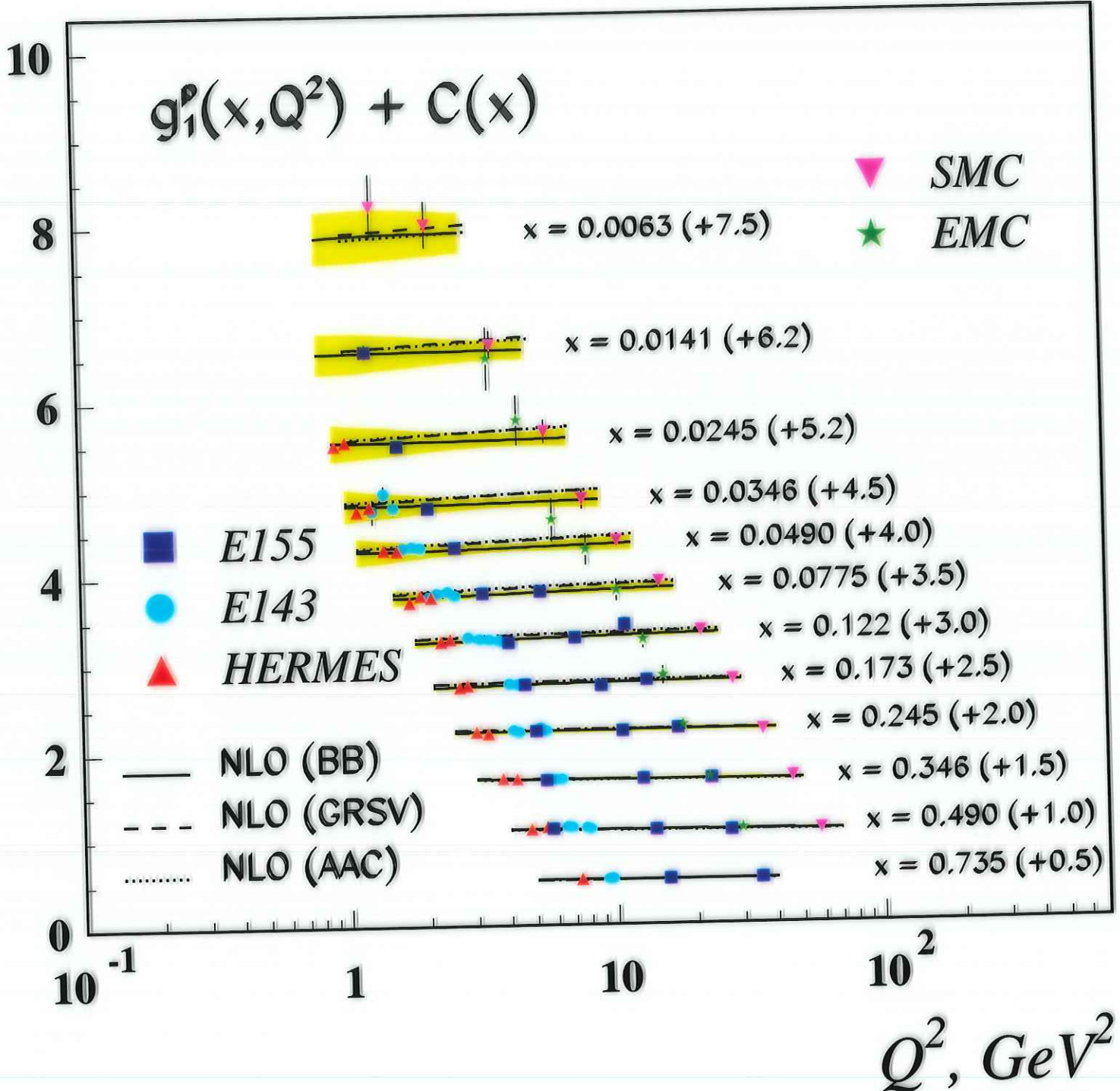


⇒ Question: Which Parameterization to be used?

⇒ Answer: R1998

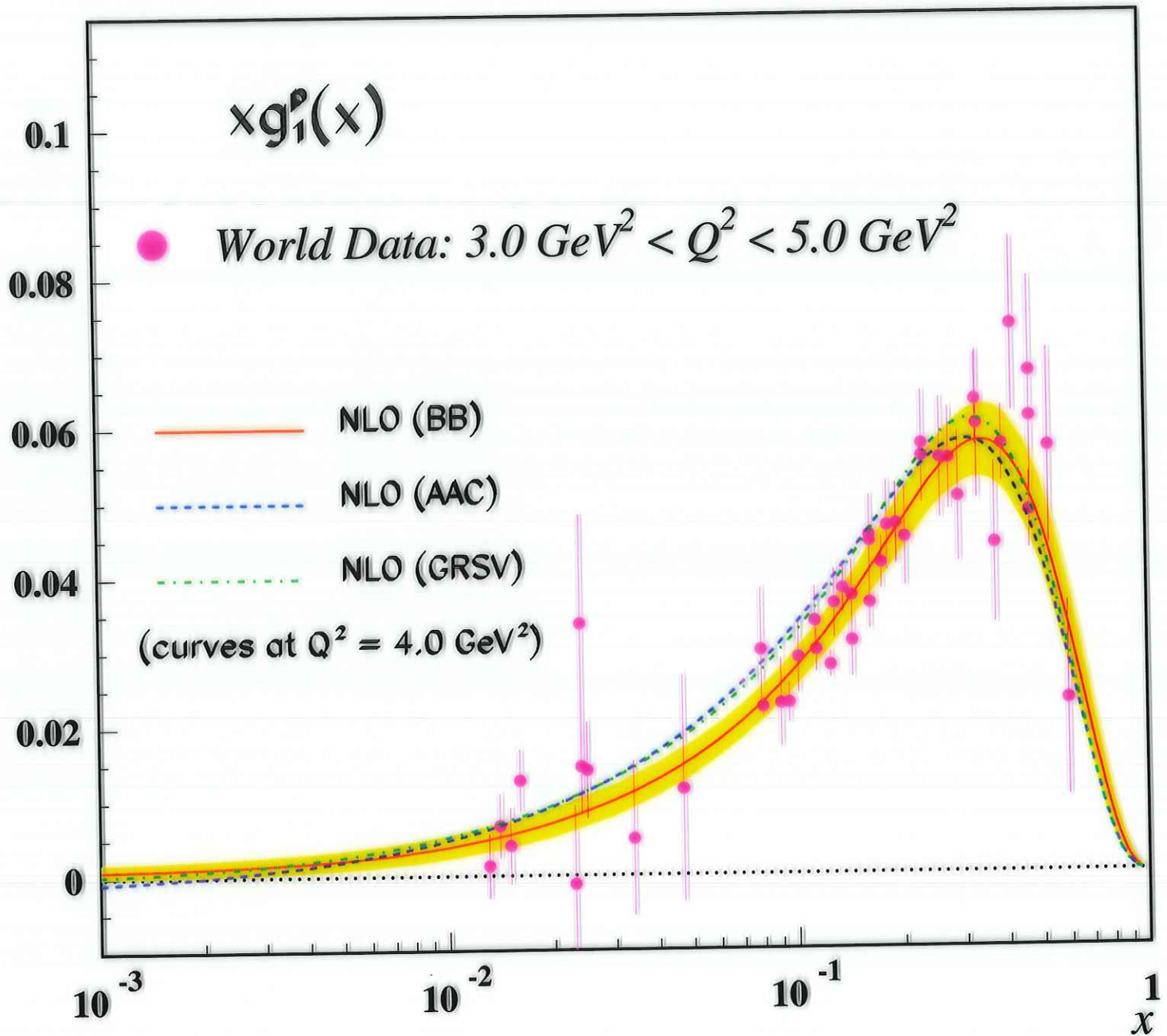


# $g_1^p(x)$ versus $Q^2$



$\Rightarrow$  Yellow error band: Fully correlated  $1\sigma$  Gaussian error propagation through the evolution equation.

# $xg_1^p(x)$ from Measured Asymmetry Data



⇒ Yellow error band: Fully correlated  $1\sigma$  Gaussian error propagation through the evolution equation.

## What about the Errors?

---

⇒ **Problem:** Systematic errors are known to be partly correlated which would lead to an overestimation of the errors when added in quadrature with the statistical ones.

- **Statistical Errors:**

To treat all data sets on the same footing statistical errors are taken only. Accept only fits with a **Positive Definite Covariance Matrix**.

⇒ Calculate the **Fully Correlated 1σ Error Bands** by Gaussian error propagation.

- **Systematic Uncertainties:**

Allow for a **Relative Normalization Shift** between the different data sets within the normalization uncertainties quoted by the experiments (**fitted and then fixed**).

$$\chi^2 = \sum_{i=1}^{n^{exp}} \left[ \frac{(N_i - 1)^2}{(\Delta N_i)^2} + \sum_{j=1}^{n^{data}} \frac{(N_i g_{1,j}^{data} - g_{1,j}^{theor})^2}{(\Delta g_{1,j}^{data})^2} \right]$$

⇒ Thereby accounting for the **main systematic uncertainties** (luminosity and beam and target polarization).



# Gaussian Error Propagation

---

In the treatment used in our analysis the evolved polarized parton densities are **linear functions of the input densities** for all parameters, except  $\Lambda_{QCD}$ .

Let  $f(x, Q^2; a_i|_{i=1}^k)$  be the evolved density at  $Q^2$  depending on the fitted parameters  $a_i|_{i=1}^k$  at the **input scale  $Q_0^2$** . Then its **fully correlated error  $\Delta f$**  as given by Gaussian error propagation is

$$\Delta f(x, Q^2) = \left[ \sum_{i=1}^k \left( \frac{\partial f}{\partial a_i} \right)^2 C(a_i, a_i) + \sum_{i \neq j=1}^k \left( \frac{\partial f}{\partial a_i} \frac{\partial f}{\partial a_j} \right) C(a_i, a_j) \right]^{\frac{1}{2}}.$$

$C(a_i, a_j)$  are the elements of the covariance matrix determined in the  $QCD$  analysis at the input scale  $Q_0^2$ .

⇒ All what is needed are the gradients  $\partial f / \partial a_i$  w.r.t. the parameters  $a_i$ . They can be calculated analytically at the input scale  $Q_0^2$ . Their value at  $Q^2$  is then given by evolution.

## Error Propagation in MELLIN-N space

---

The general form of the derivative of the MELLIN moment  $\mathbf{M}[f(a)](N)$  w.r.t. parameter  $a$  for complex values of  $N$  is

$$\frac{\partial \mathbf{M}[f(a)](N)}{\partial a} = F(a) \times \mathbf{M}[f(a)](N),$$

- For  $\Delta u_v$  and  $\Delta d_v$ :

$$F(a_i) = \frac{\psi(N - 1 + a_i) - \psi(N + a_i + b_i) + \gamma_i(b_i + 1)}{(N + a_i + b_i)(N + a_i + b_i + \gamma_i(N - 1 + a_i))} - \frac{\psi(a_i) + \psi(a_i + b_i + 1) - \gamma_i(b_i + 1)}{(a_i + b_i + 1)(a_i + b_i + 1 + \gamma_i a_i)},$$

$$F(b_i) = \frac{\psi(b_i + 1) - \psi(N + a_i + b_i) - \gamma_i(N - 1 + a_i)}{(N + a_i + b_i)(N + a_i + b_i + \gamma_i(N - 1 + a_i))} - \frac{\psi(b_i + 1) + \psi(a_i + b_i + 1) + \gamma_i a_i}{(a_i + b_i + 1)(a_i + b_i + 1 + \gamma_i a_i)}$$

Note:

$$x \Delta q_i(x, Q_0^2) = \eta_i A_i x^{a_i} (1 - x)^{b_i} (1 + \gamma_i x)$$

## Error Propagation in MELLIN-N space (cont'd)

---

- For  $\Delta\bar{q}$  and  $\Delta G$ :

$$F(\eta_i) = \frac{1}{\eta_i},$$

$$F(a_i) = \psi(N - 1 + a_i) - \psi(N + a_i + b_i) \\ - \psi(a_i) + \psi(a_i + b_i + 1).$$

Note:

$$x\Delta q_i(x, Q_0^2) = \eta_i A_i x^{a_i} (1-x)^{b_i}$$

with  $\psi(z) = d/dz(\log\Gamma(z))$  the EULER  $\psi$ -function.

---

⇒ The gradients evolved in **MELLIN-N space** are then transformed back to  **$x$ -space** and can be used according to the error propagation equation.

---

- When fitting  $\Lambda_{QCD}$  its gradient has to be determined numerically due to non-linear and iterative aspects in the calculation of  $\alpha_s(Q^2, \Lambda_{QCD})$ :

$$\frac{\partial f(x, Q^2, \Lambda)}{\partial \Lambda} = \frac{f(x, Q^2, \Lambda + \delta) - f(x, Q^2, \Lambda - \delta)}{2\delta}$$

with  $\delta \sim 10 \text{ MeV}$ .



## Parameter Values at $Q_0^2 = 4.0 \text{ GeV}^2$

---

7+1 Parameter Fit based on the Asymmetry Data:

	Scenario 1			
	LO		NLO	
	value	error	value	error
$\Lambda_{QCD}^{(4)}, \text{ MeV}$	203	120	235	53
$\eta_{uv}$	0.926	fixed	0.926	fixed
$a_{uv}$	0.197	0.013	0.294	0.035
$b_{uv}$	2.403	0.107	3.167	0.212
$\gamma_{uv} (*)$	21.34	fixed	27.22	fixed
$\eta_{dv}$	-0.341	fixed	-0.341	fixed
$a_{dv}$	0.190	0.049	0.254	0.111
$b_{dv}$	3.240	0.884	3.420	1.332
$\gamma_{dv} (*)$	30.80	fixed	19.06	fixed
$\eta_{sea}$	-0.353	0.054	-0.447	0.082
$a_{sea}$	0.367	0.048	0.424	0.062
$b_{sea} (*)$	8.51	fixed	8.93	fixed
$\eta_G$	1.281	0.816	1.026	0.554
$a_G$	$a_{sea} + 0.9$		$a_{sea} + 1.0$	
$b_G (*)$	5.91	fixed	5.51	fixed
$\chi^2 / \text{NDF}$	1.02		0.90	

⇒ The parameters marked by (\*) have been fitted first and then fixed since the present data do not constrain their values well enough.

⇒ Scenario 2 :  $a_G = a_{sea} + 0.6$  (LO)  
 $a_G = a_{sea} + 0.5$  (NLO)

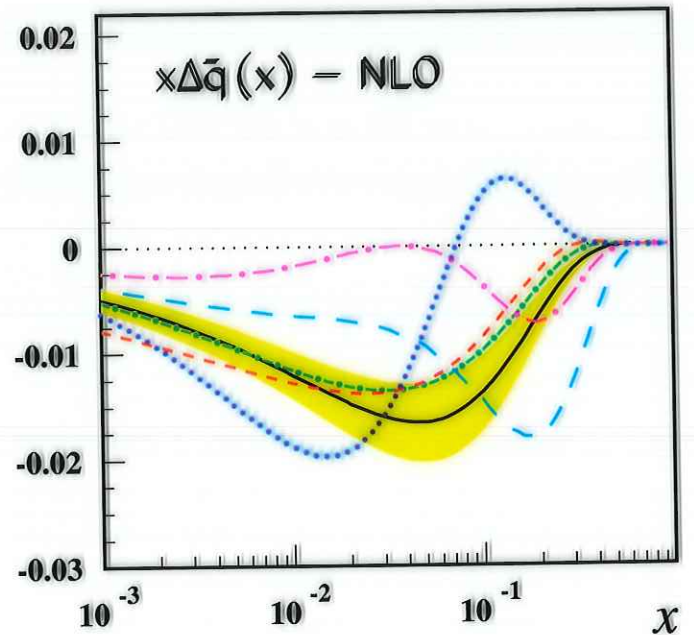
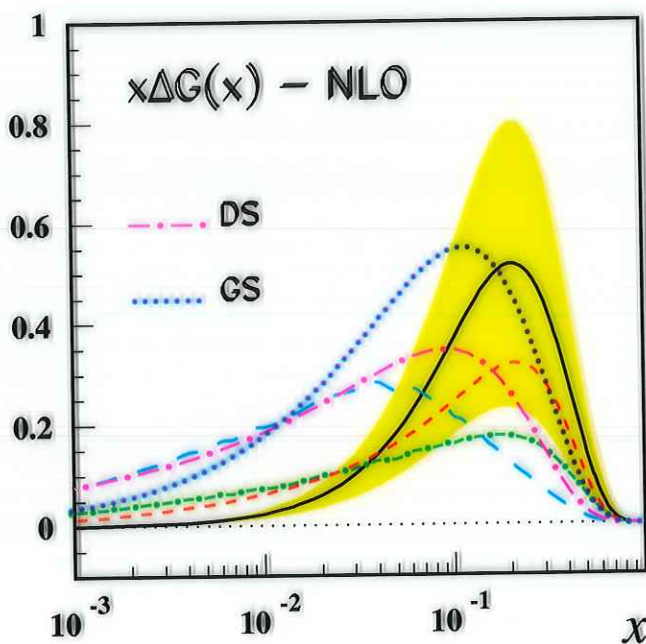
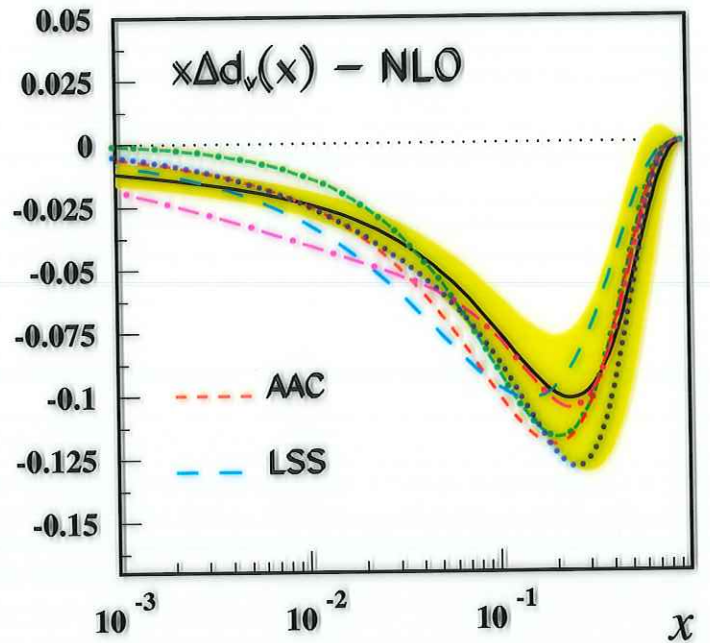
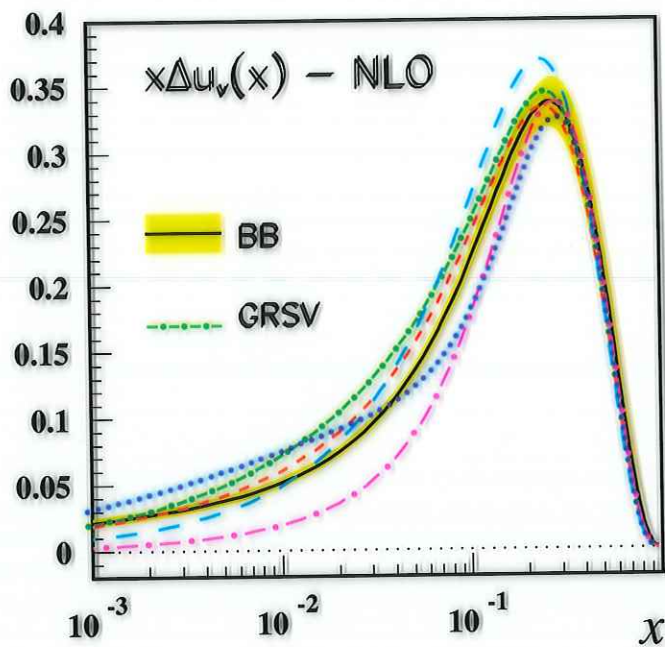
Covariance Matrices at  $Q_0^2 = 4.0 \text{ GeV}^2 - 7 + 1$  Parameter Fit - Scenario 1

LO									
	$\Lambda_{QCD}^{(4)}$	$a_{uv}$	$b_{uv}$	$a_{dv}$	$b_{dv}$	$\eta_{sea}$	$a_{sea}$	$\eta_G$	
$\Lambda_{QCD}^{(4)}$	1.43E-2								
$a_{uv}$	-2.05E-5	1.80E-4							
$b_{uv}$	-9.07E-5	3.91E-4	1.15E-2						
$a_{dv}$	1.10E-4	1.03E-5	-2.40E-3	2.43E-3					
$b_{dv}$	-4.65E-5	-7.92E-3	-6.86E-3	5.48E-3	7.82E-01				
$\eta_{sea}$	1.02E-4	-4.46E-4	-2.84E-3	9.85E-4	2.82E-2	2.94E-3			
$a_{sea}$	-4.31E-5	1.58E-4	1.33E-3	-5.96E-4	-9.32E-3	-2.58E-4	2.29E-3		
$\eta_G$	-1.03E-3	2.02E-3	1.58E-2	-2.78E-3	-1.61E-1	-1.59E-2	9.56E-3	6.65E-1	

NLO									
	$\Lambda_{QCD}^{(4)}$	$a_{uv}$	$b_{uv}$	$a_{dv}$	$b_{dv}$	$\eta_{sea}$	$a_{sea}$	$\eta_G$	
$\Lambda_{QCD}^{(4)}$	2.81E-3								
$a_{uv}$	2.71E-5	1.22E-3							
$b_{uv}$	-1.30E-4	5.10E-3	4.50E-2						
$a_{dv}$	-3.35E-4	-5.17E-4	-3.23E-3	1.23E-2					
$b_{dv}$	-6.22E-4	-1.27E-2	4.65E-2	8.29E-2	1.78E-0				
$\eta_{sea}$	-5.30E-5	-2.13E-3	-1.12E-2	5.19E-3	4.74E-2	6.77E-3			
$a_{sea}$	-4.85E-6	9.07E-4	4.49E-3	-3.78E-3	-2.98E-2	-2.39E-3	3.82E-3		
$\eta_G$	4.03E-4	1.41E-2	6.71E-2	-3.07E-2	-2.22E-1	-3.78E-2	1.90E-2	3.07E-1	



# Pol. Parton Densities at $Q^2 = 4.0 \text{ GeV}^2$

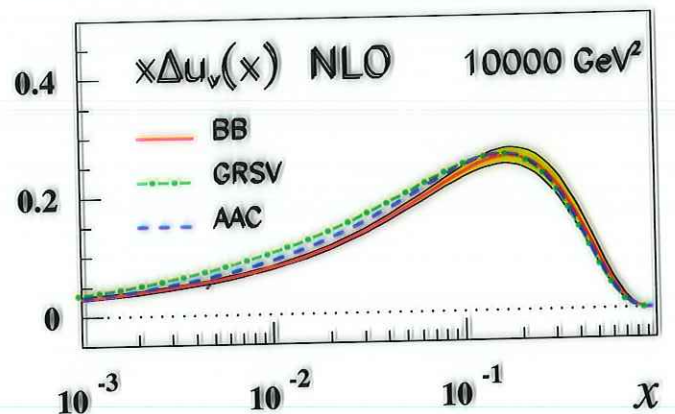
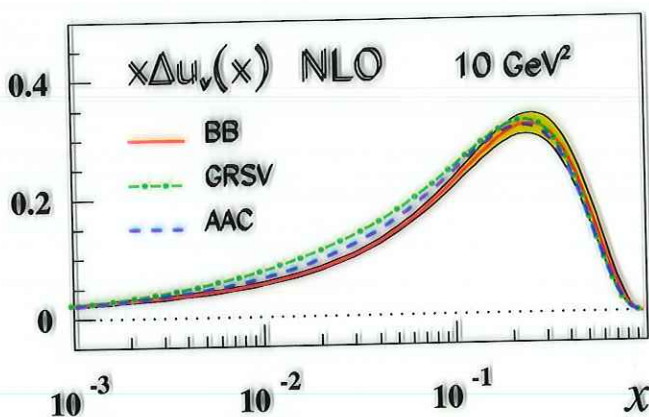
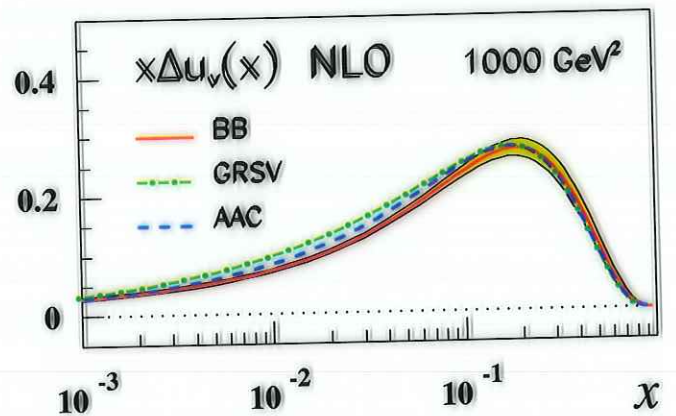
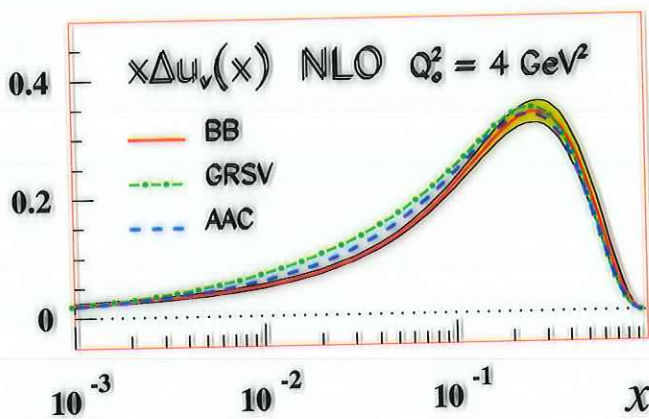
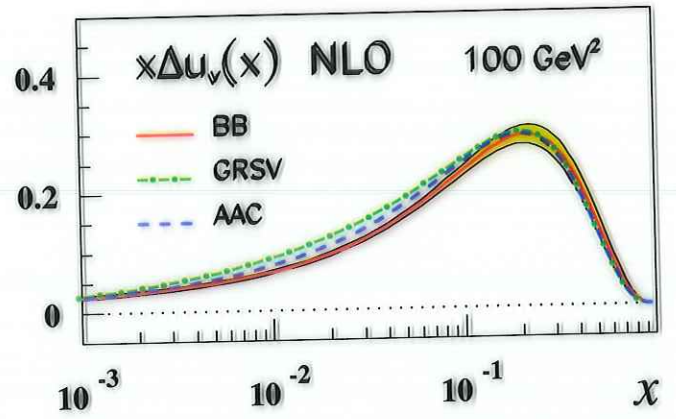
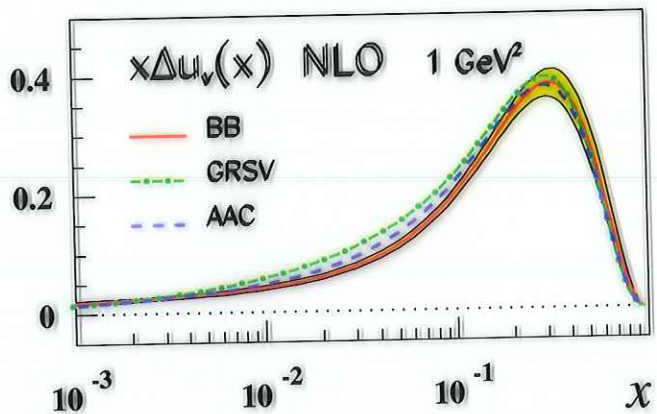


⇒ Yellow band: Fully correlated  $1\sigma$  statistical error band from the BB analysis.



# Evolution of Polarized Parton Densities

- 7+1 Parameter Fit based on the Asymmetry Data:

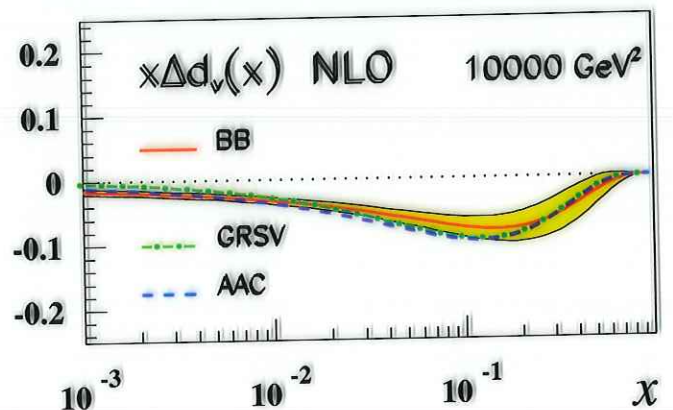
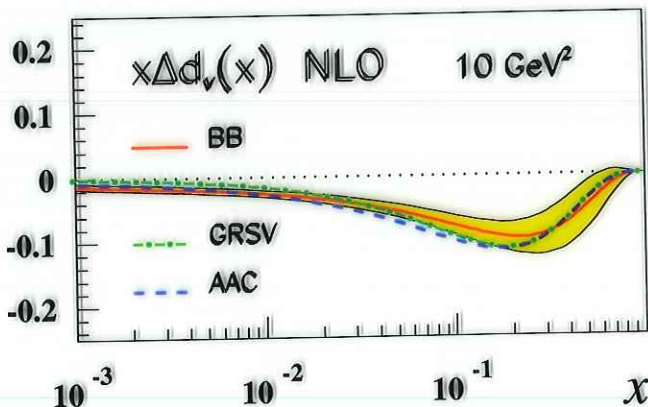
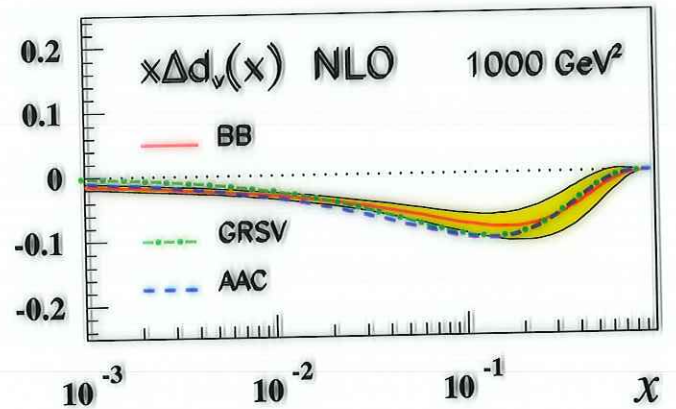
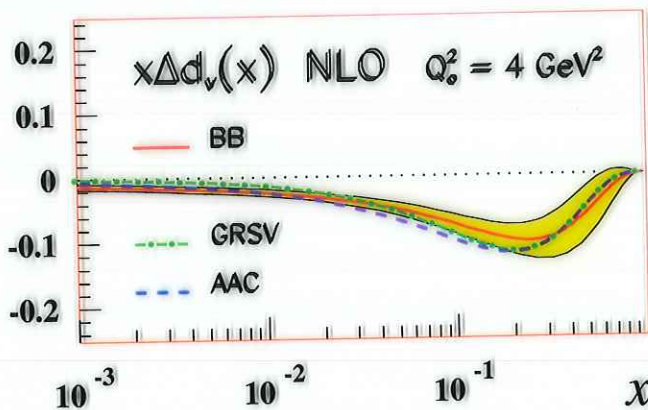
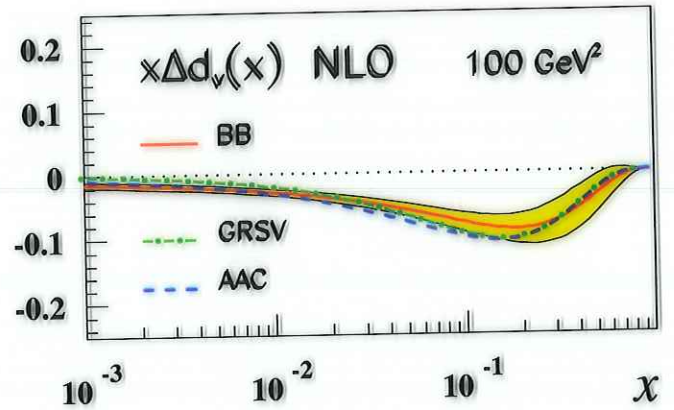
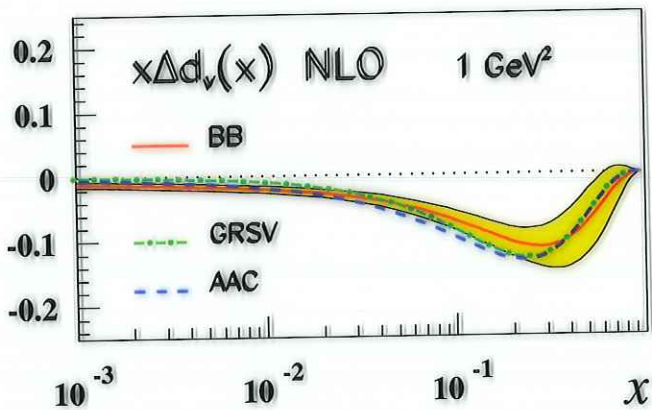


⇒ Yellow error band:

Fully correlated  $1\sigma$  Gaussian error propagation through the evolution equation.

# Evolution of Polarized Parton Densities

- 7+1 Parameter Fit based on the Asymmetry Data:

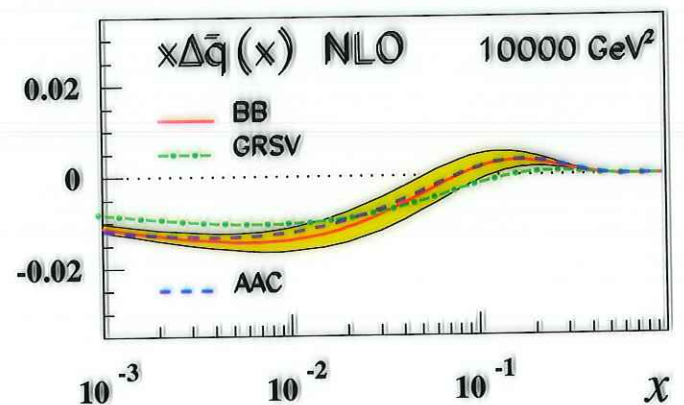
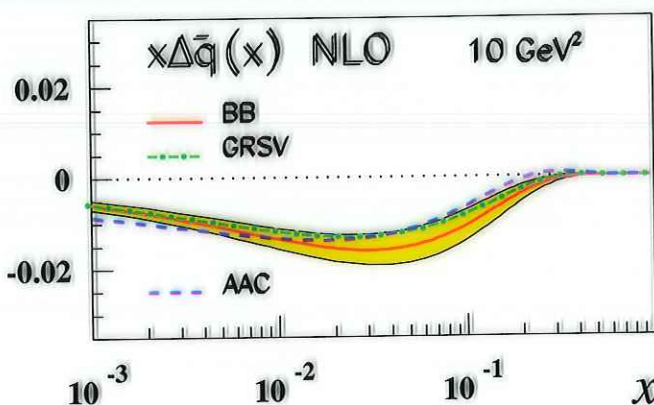
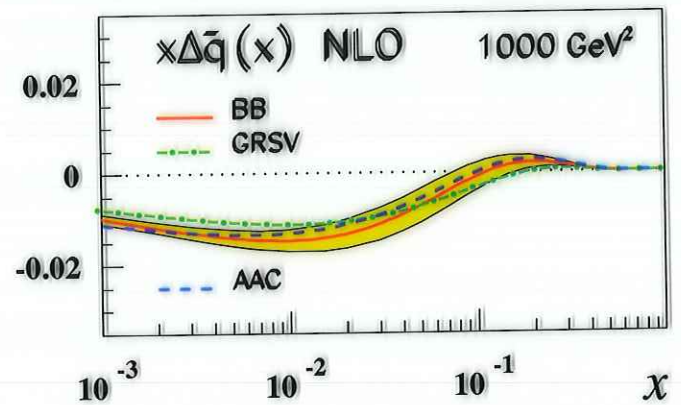
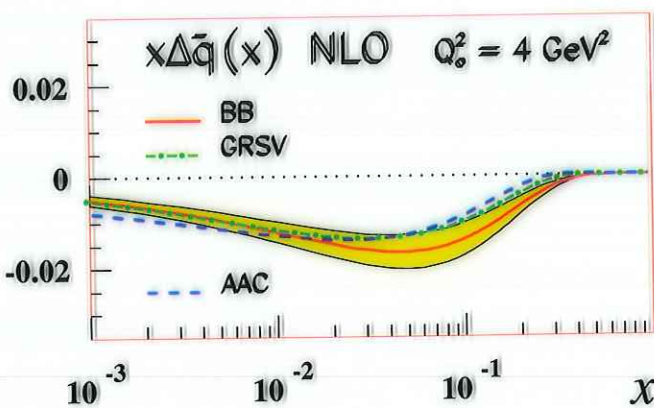
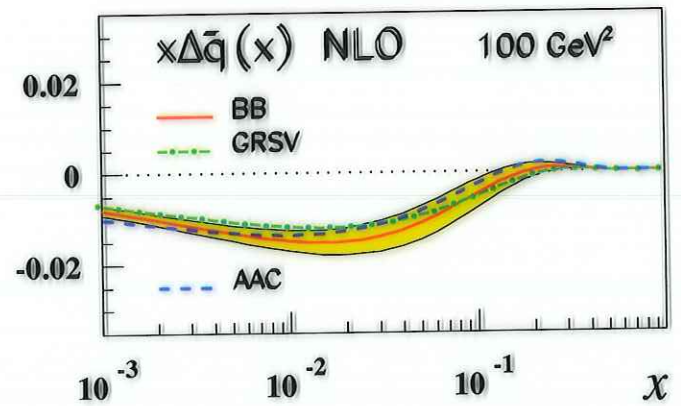
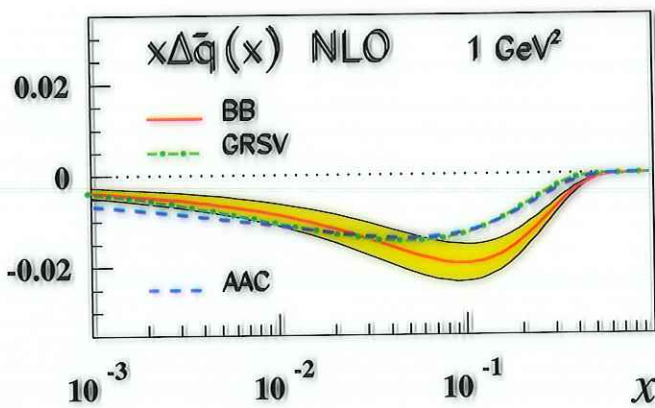


⇒ Yellow error band: Fully correlated  $1\sigma$  Gaussian error propagation through the evolution equation.



# Evolution of Polarized Parton Densities

- 7+1 Parameter Fit based on the Asymmetry Data:



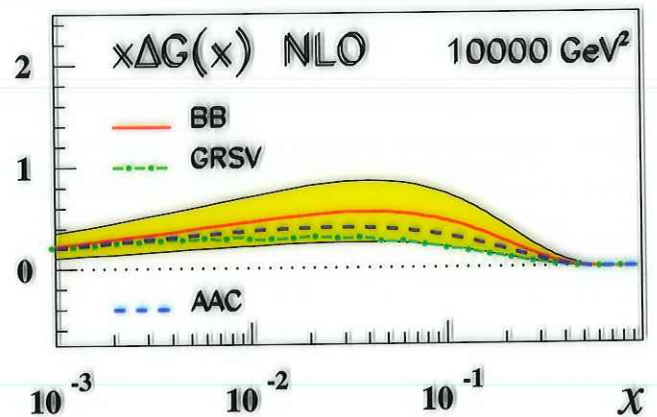
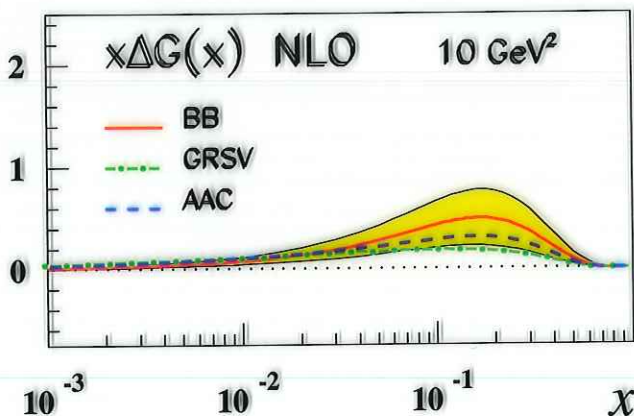
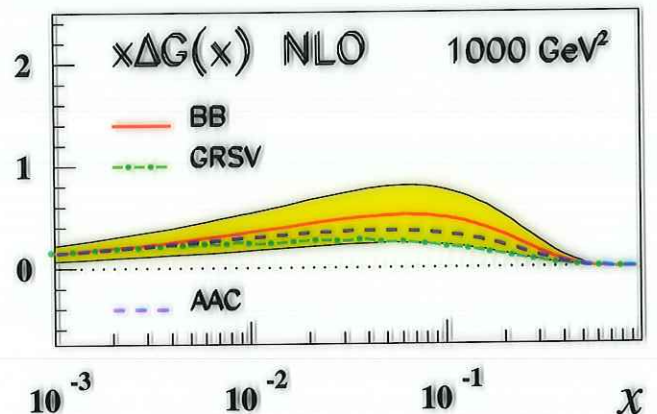
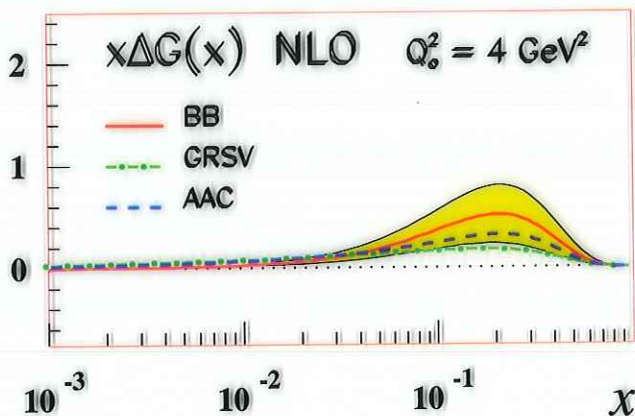
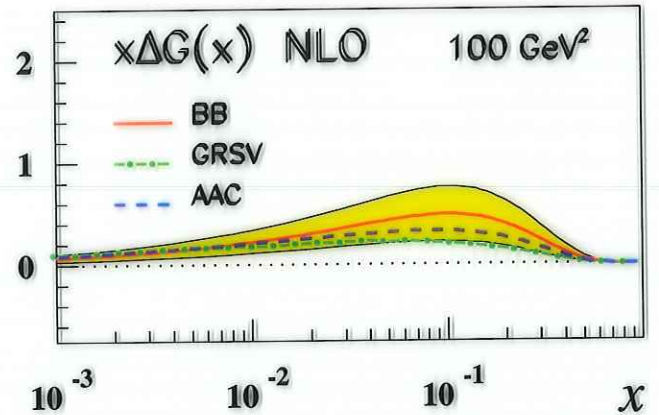
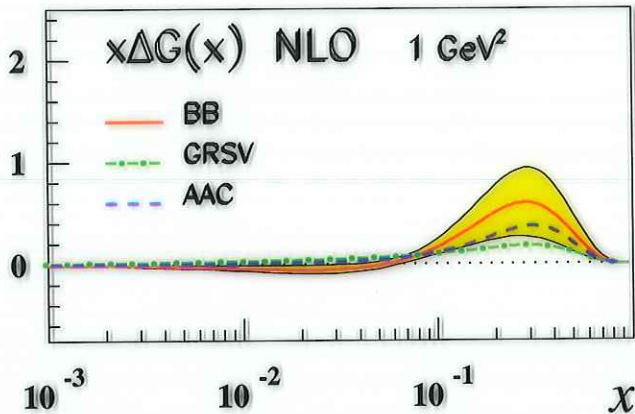
⇒ Yellow error band:

Fully correlated  $1\sigma$  Gaussian error propagation through the evolution equation.



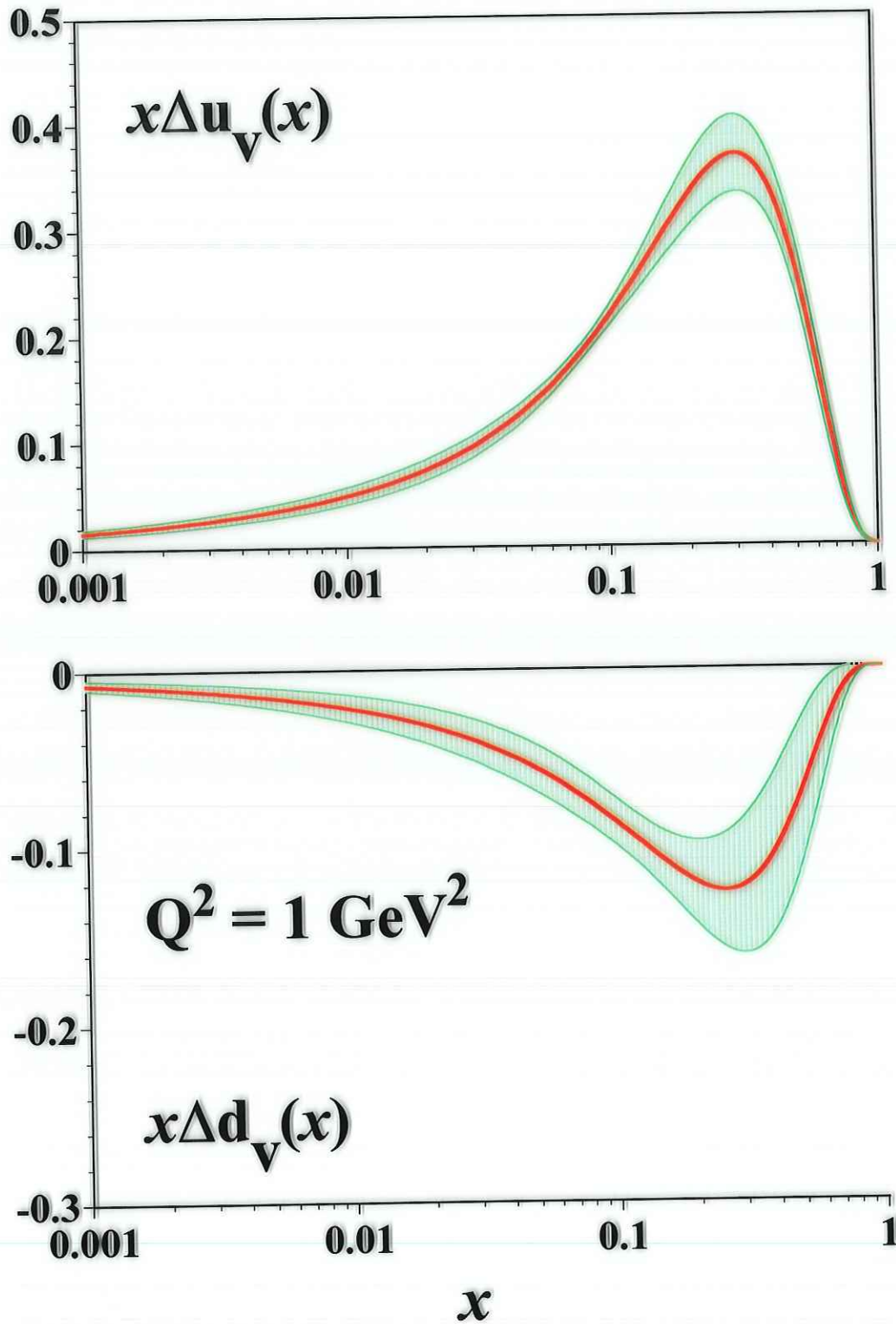
# Evolution of Polarized Parton Densities

- 7+1 Parameter Fit based on the Asymmetry Data:



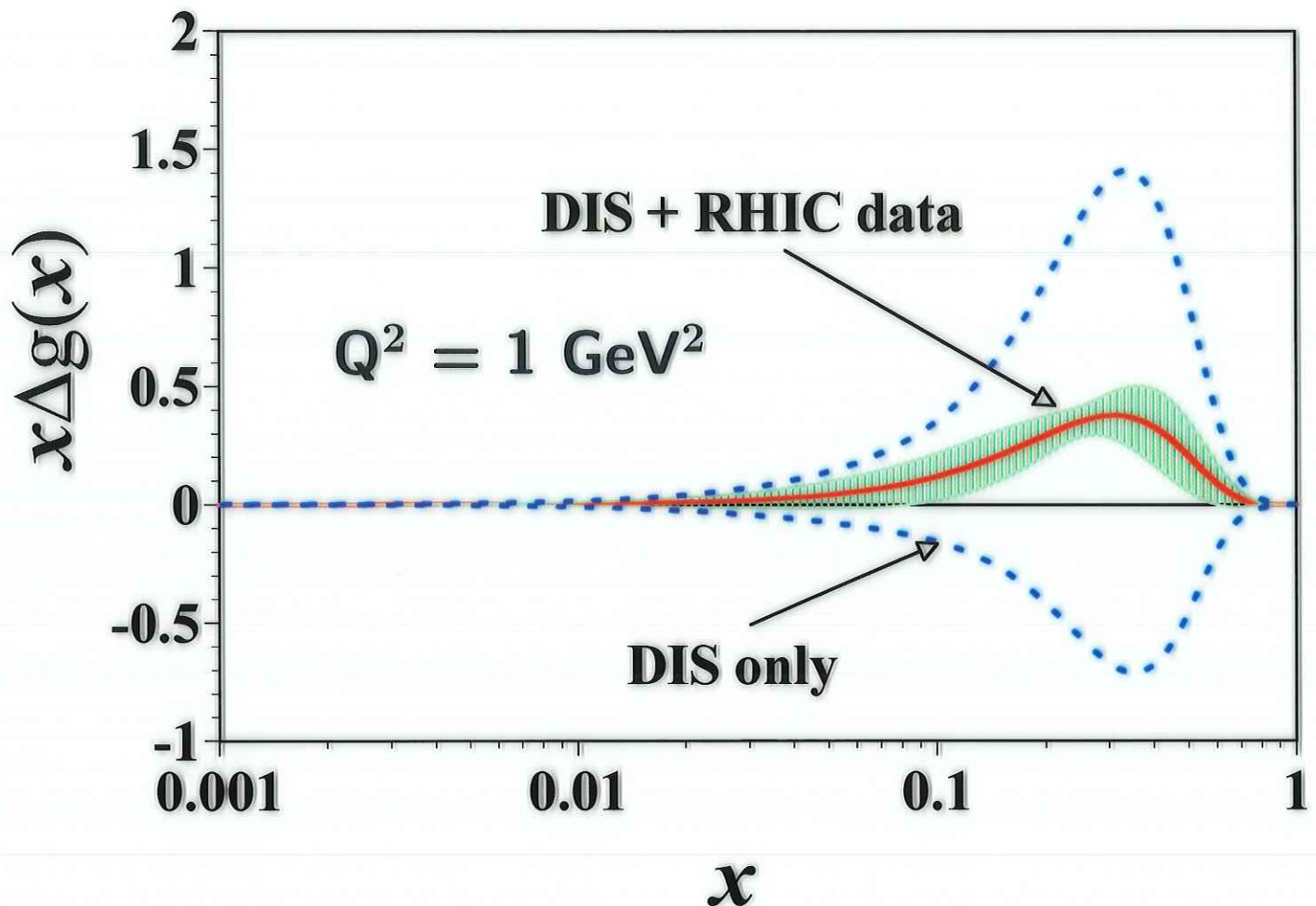
⇒ Yellow error band: Fully correlated  $1\sigma$  Gaussian error propagation through the evolution equation.

# AAC: Polarized Valence Quark Densities



⇒ Green band: Fully correlated  $1\sigma$  statistical error band.

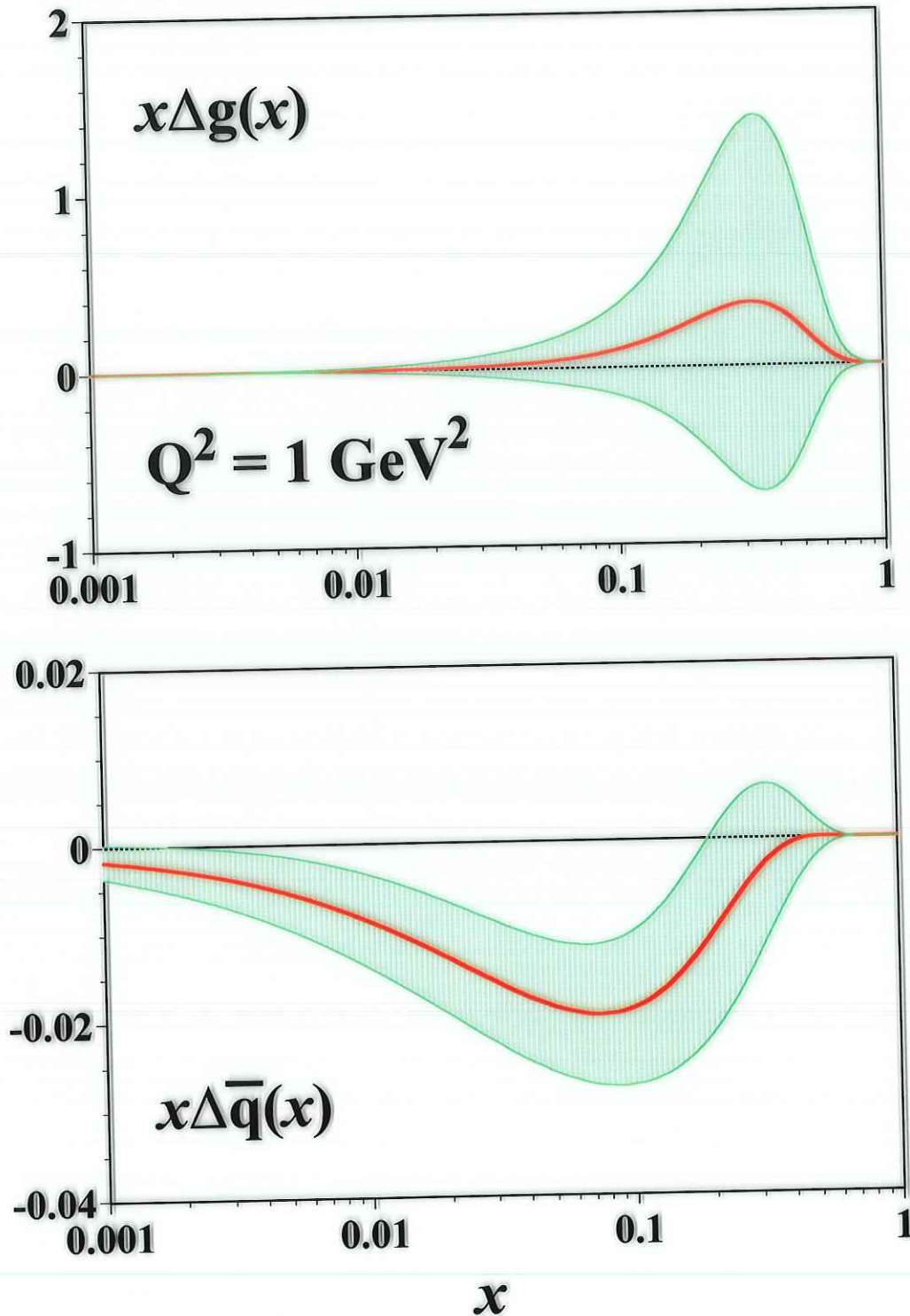
# AAC: Polarized Gluon Density with RHIC



⇒ Green band: Fully correlated  $1\sigma$  statistical error band.



# AAC: Polarized Gluon and Sea Densities

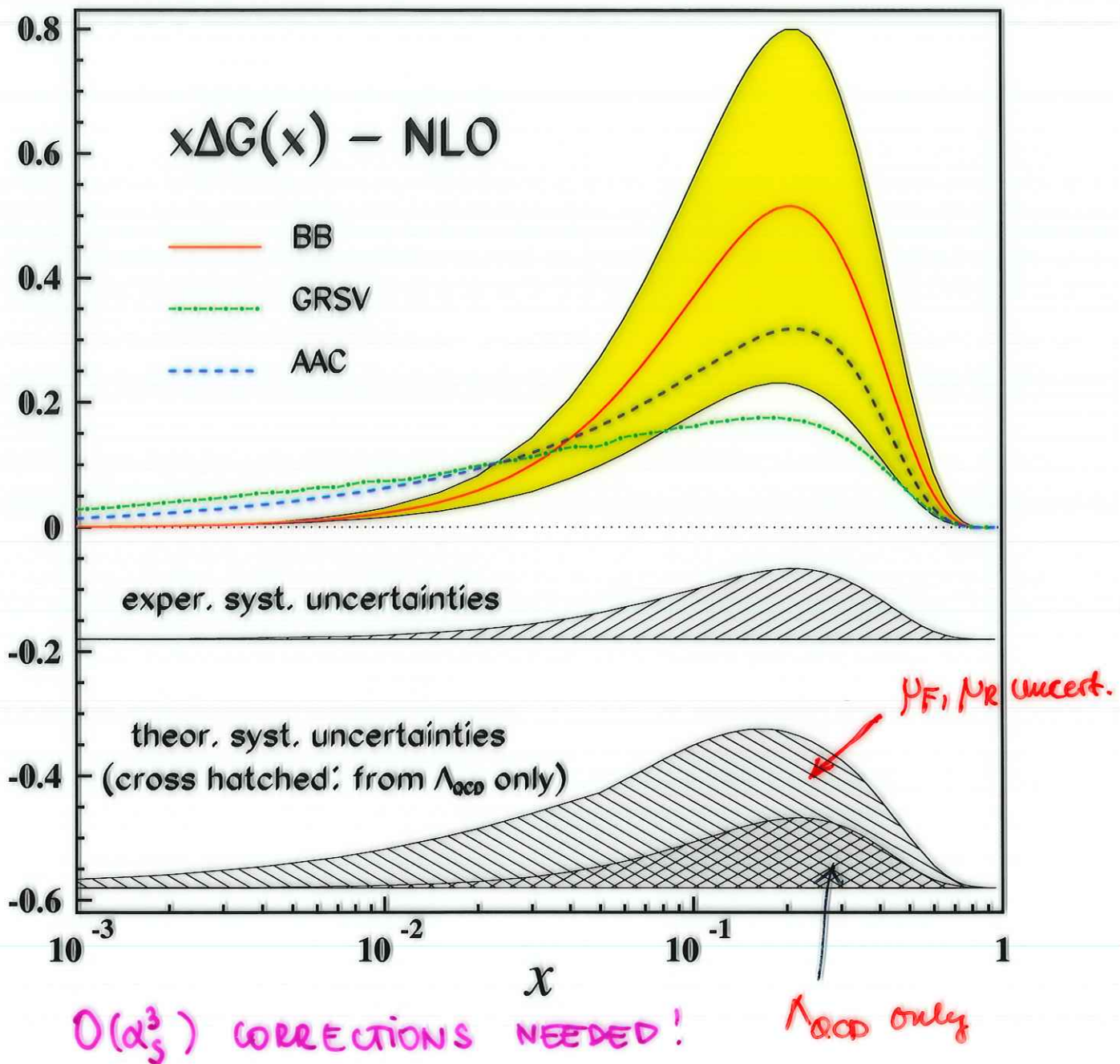


⇒ Green band: Fully correlated  $1\sigma$  statistical error band.

# Theoretical Systematics for the Gluon

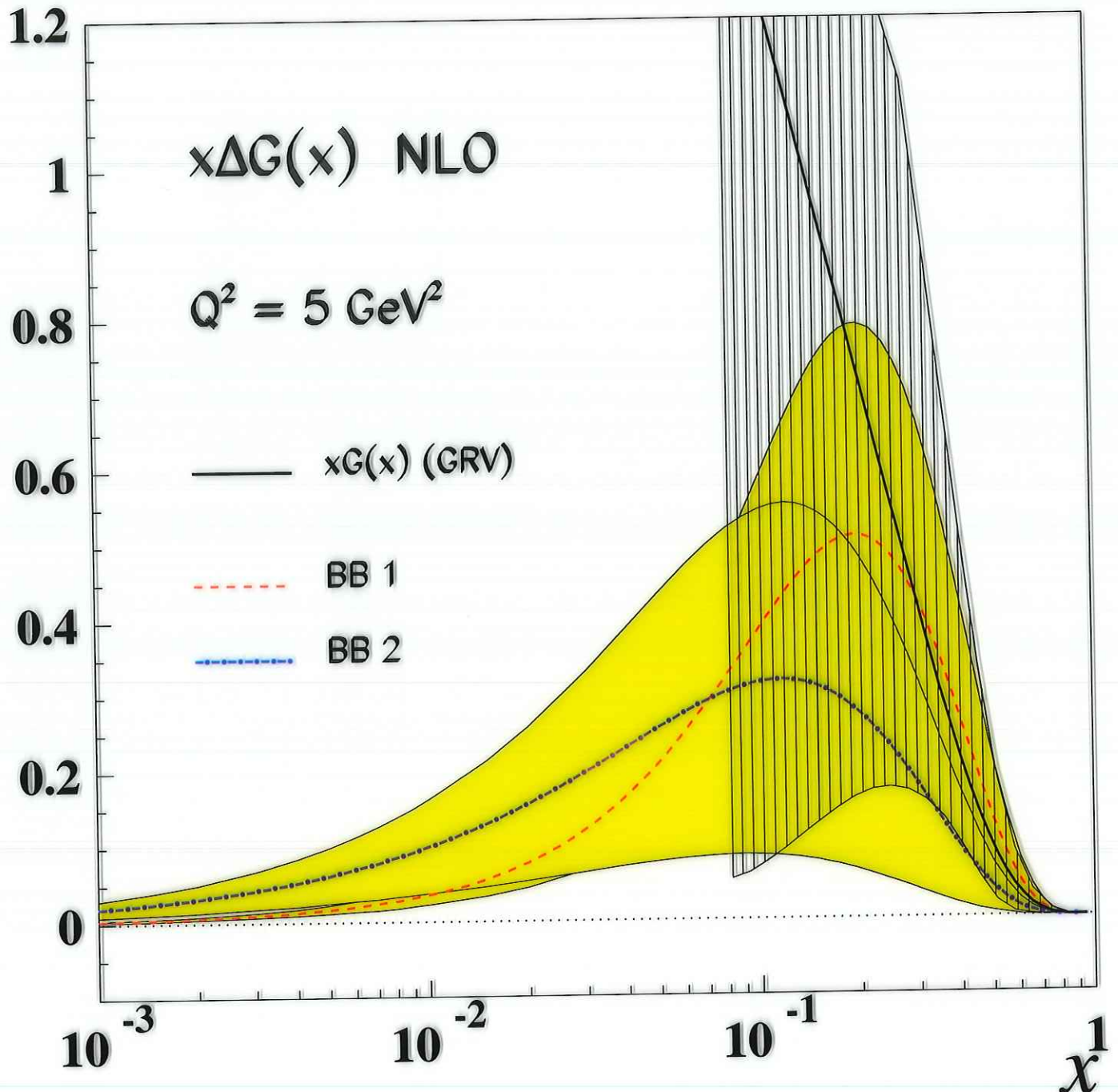
- 7+1 Parameter Fit based on the Asymmetry Data:

⇒ Scenario 1



# The Polarized Gluon at $Q_0^2 = 5.0 \text{ GeV}^2$

- 7+1 Parameter Fit based on the Asymmetry Data:

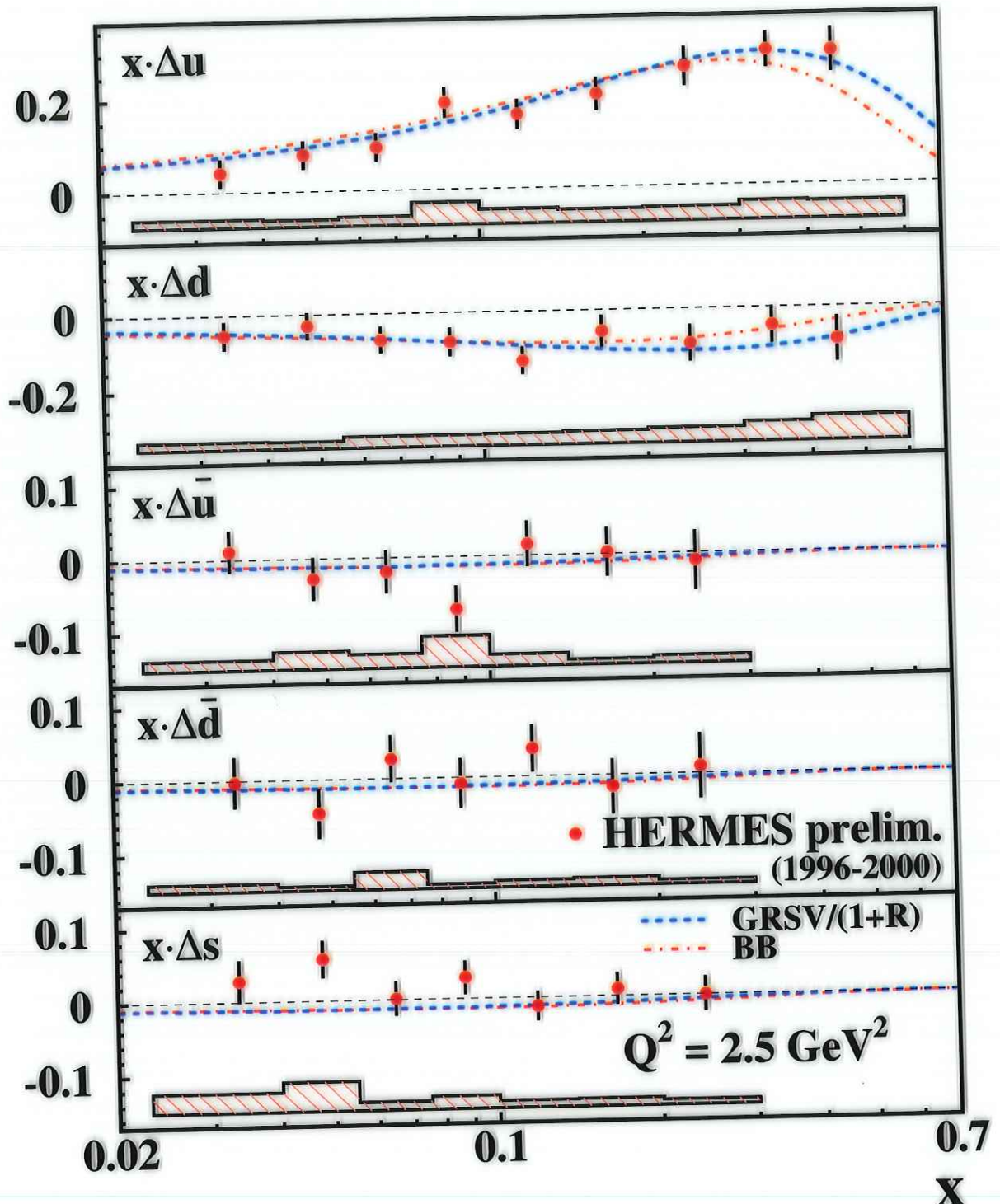


⇒ Yellow error bands: Fully correlated  $1\sigma$  Gaussian error propagation at  $Q^2 = 5.0 \text{ GeV}^2$ .

⇒ Hatched Area: Error Band taken from H1 and laid over the GRV curve.

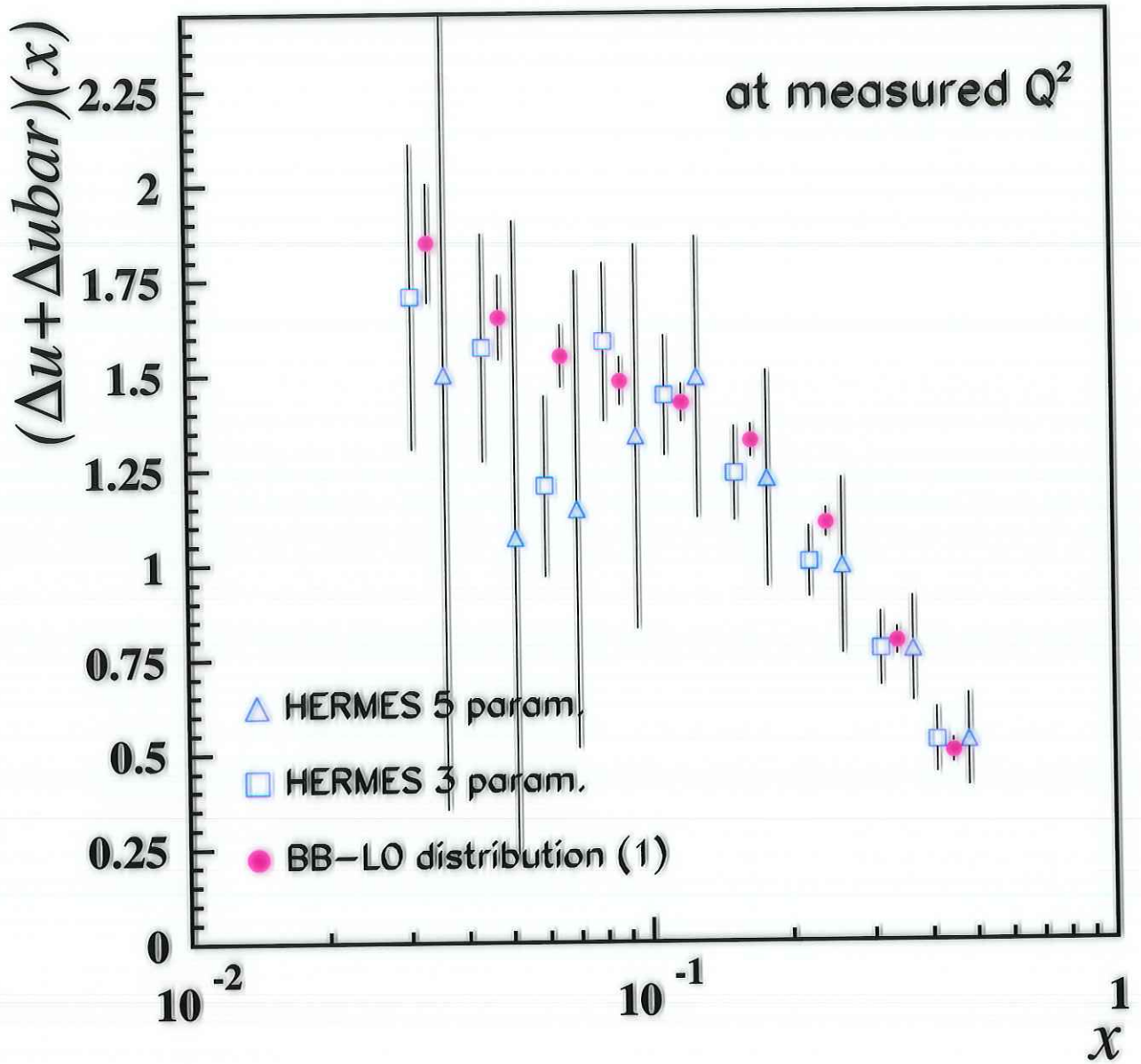


# Extraction of $\Delta q$ from Semi-Incl. Data



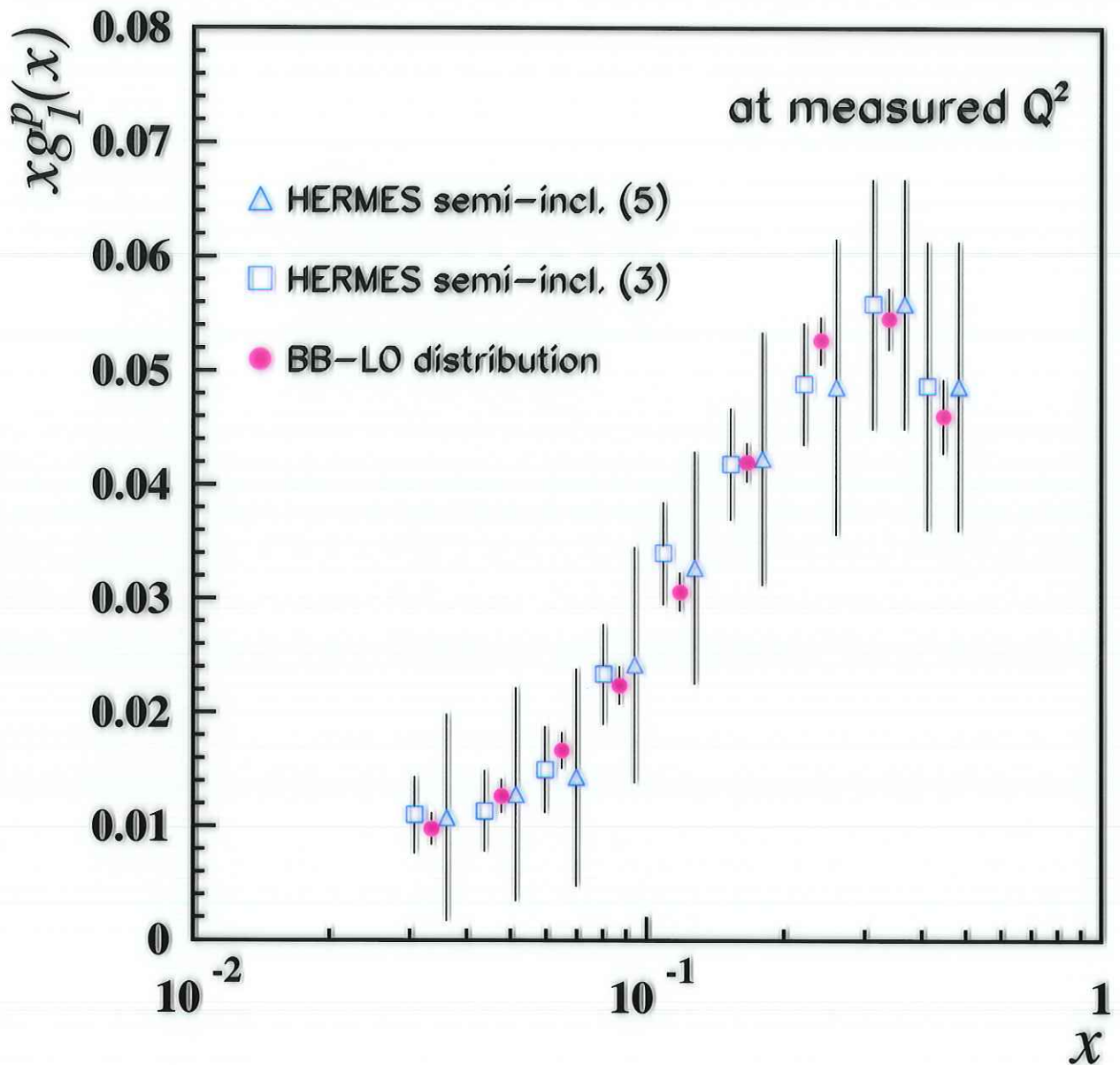
- $z$ -range in Semi-Incl. Analysis:  $0.2 < z < 0.7$
- Symmetric Sea assumed in both the Semi-Inclusive Analysis and the LO QCD Analyses.

## Comparison with $\Delta q$ from Semi-Incl. Data



⇒  $z$ -range in Semi-Incl. Analysis:  $0.2 < z < 0.7$

## Comp. with $xg_1^p(x)$ from Semi-Incl. Data



$\Rightarrow$   $z$ -range in Semi-Incl. Analysis:  $0.2 < z < 0.7$



## SUM RULES AND INTEGRAL RELATIONS:

TWIST 2:

$$g_2(x, Q^2) = -g_1(x, Q^2) + \int_x^1 \frac{dy}{y} g_1(y, Q^2)$$

Wandzura, Wilczek, 1977;

Piccione, Ridolfi 1998; J.B., A. Tkabladze, 1998 : with TM

$$g_3(x, Q^2) = 2x \int_x^1 \frac{dy}{y^2} g_4(y, Q^2)$$

J.B., N. Kochelev, 1996; J.B., A. Tkabladze, 1998 : with TM

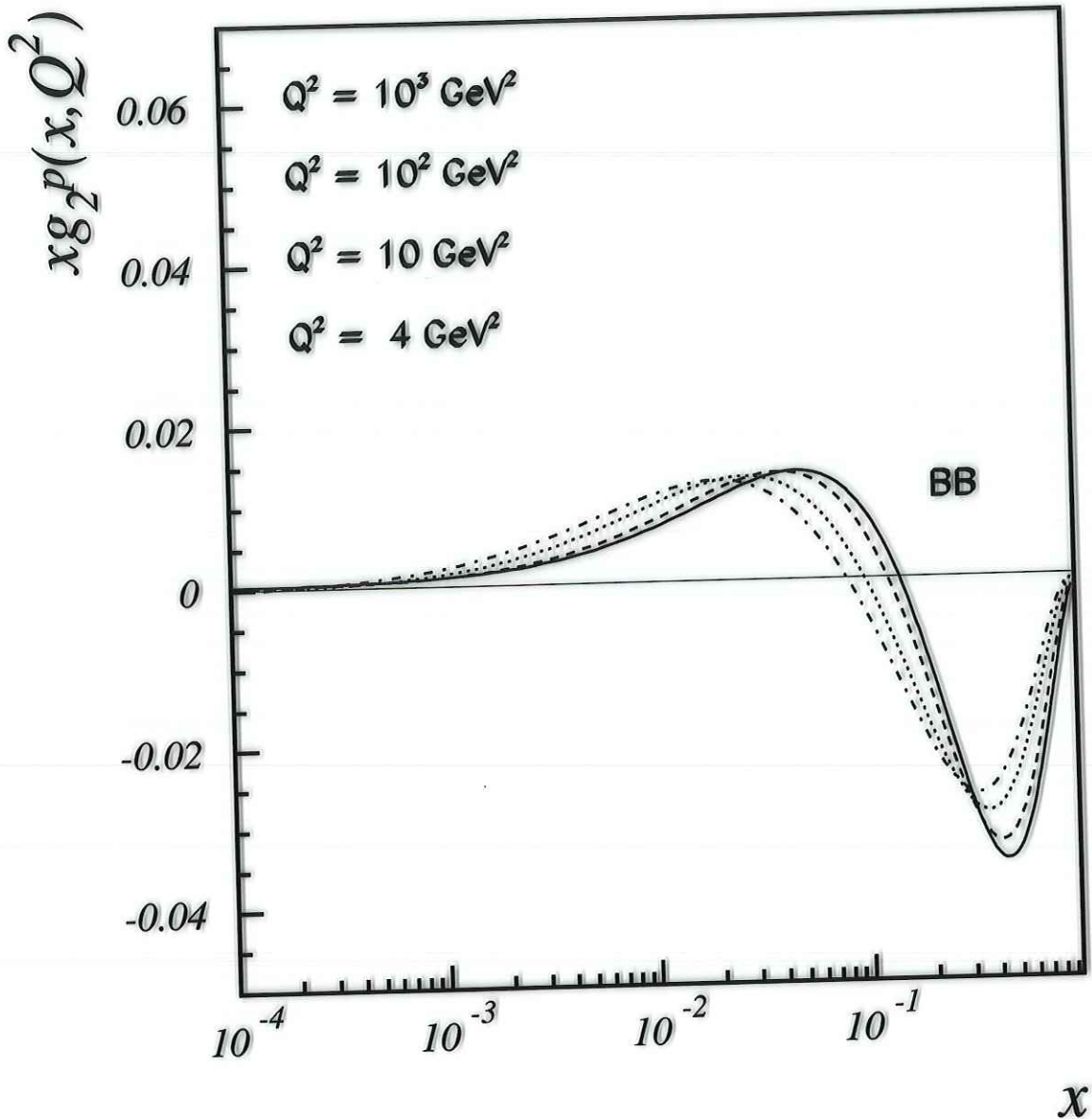
TWIST 3:

INCLUDE NUCLEON MASS EFFECTS.

J.B., A. Tkabladze, 1998

$$\begin{aligned} g_1(x, Q^2) &= \frac{4M^2 x^2}{Q^2} \left[ g_2(x, Q^2) - 2 \int_x^1 \frac{dy}{y} g_2(y, Q^2) \right] \\ \frac{4M^2 x^2}{Q^2} g_3(x, Q^2) &= g_4(x, Q^2) \left( 1 + \frac{4M^2 x^2}{Q^2} \right) + 3 \int_x^1 \frac{dy}{y} g_4(y, Q^2) \\ 2x g_5(x, Q^2) &= - \int_x^1 \frac{dy}{y} g_4(y, Q^2) \end{aligned}$$

# $g_2^{\text{light}}(x, Q^2)$



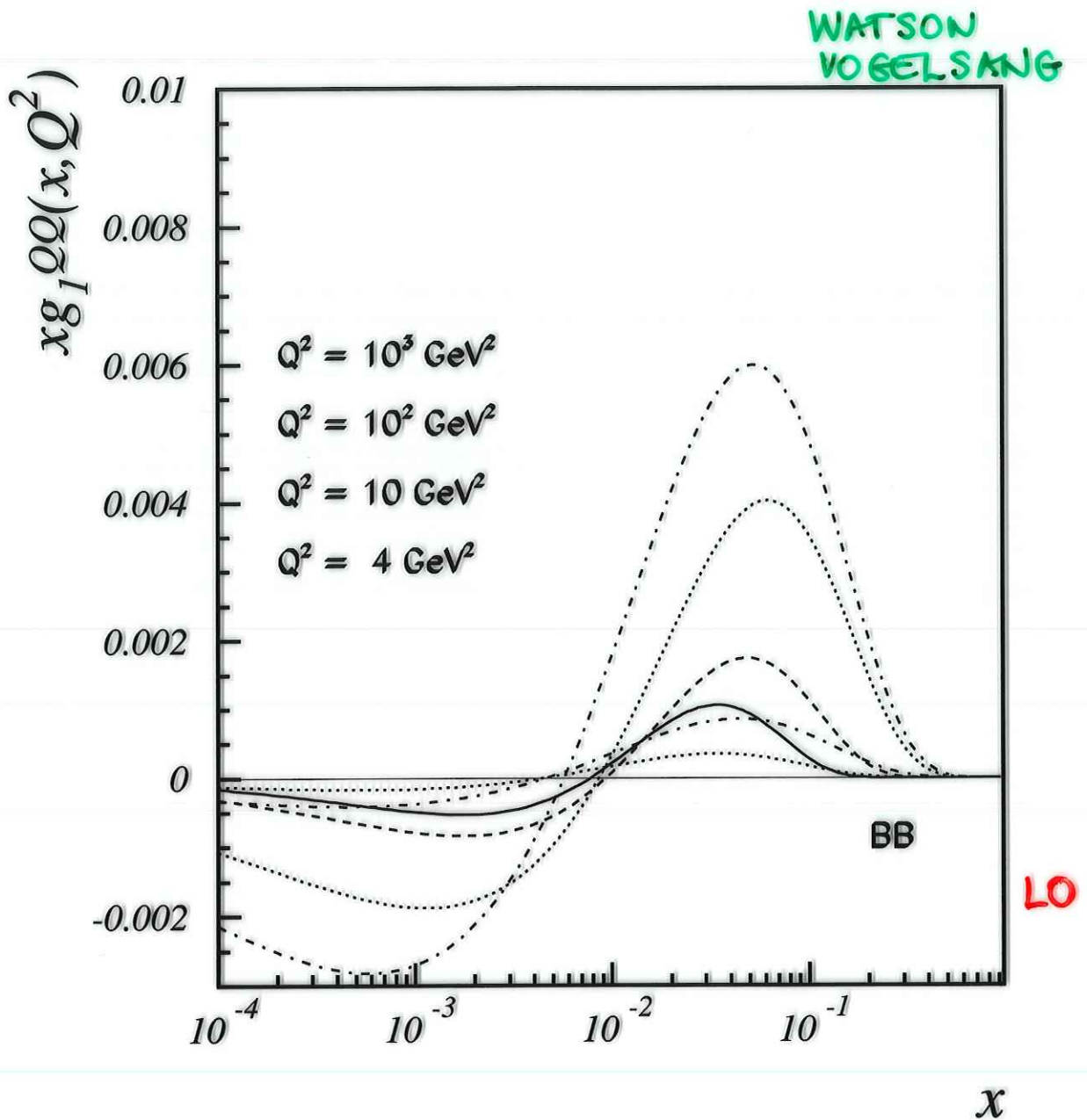
$$g_2^{\text{II}}(x, Q^2) = -g_1^{\text{II}}(x, Q^2) + \int_x^1 \frac{dz}{z} g_1^{\text{II}}(z, Q^2)$$

WW - RELATION

$$\int_0^1 dz g_2^{\text{II}}(z, Q^2) = 0 \quad \text{BC - RELATION}$$

$$g_1^{Q\bar{Q}}(x, Q^2)$$

J.B., V. RAVINDRAN, AND W.L. VAN NEERVEN, hep-ph/0304292,  
 PHYS. REV. D TO APPEAR.



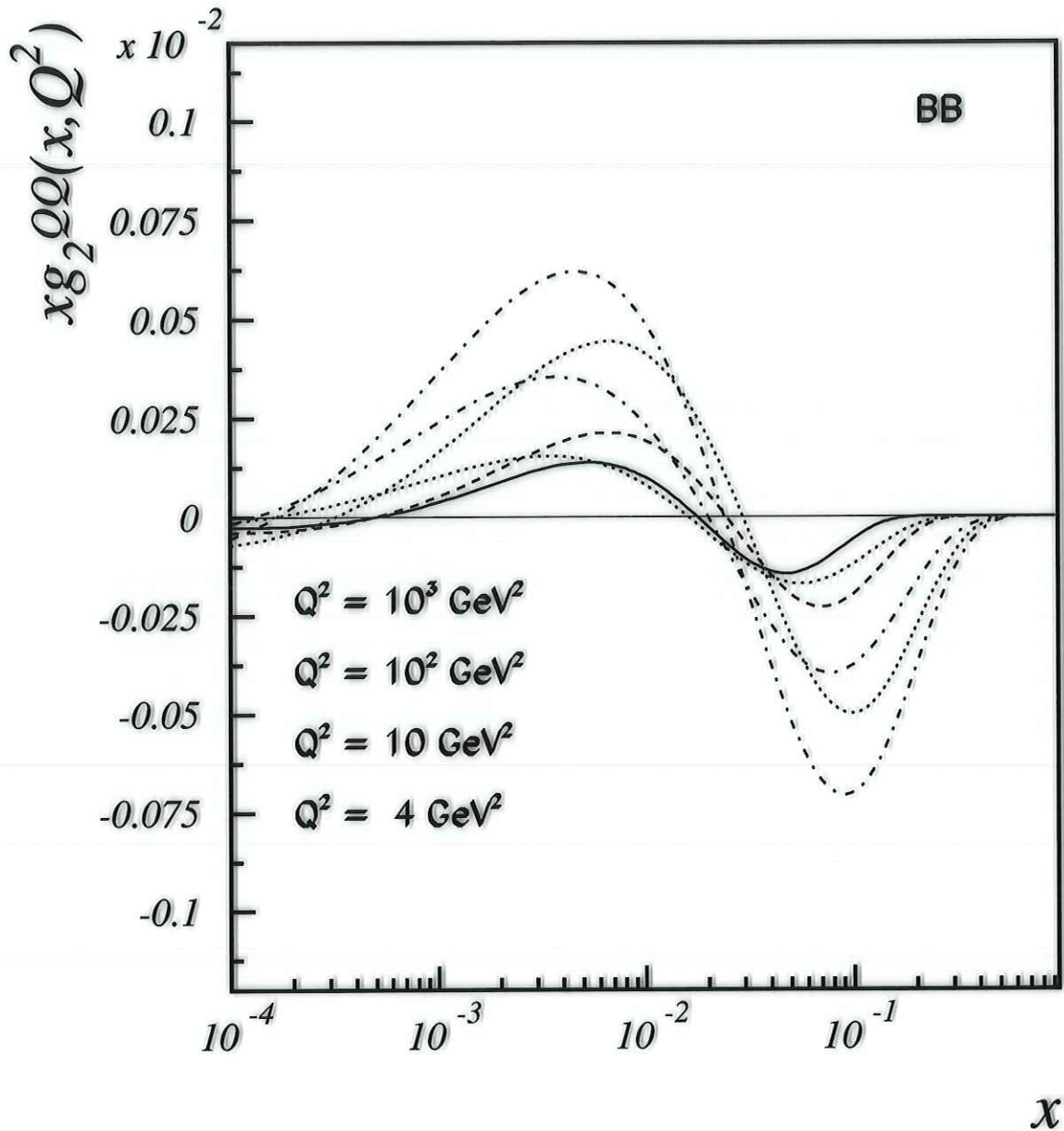
UPPER LINES:  $c\bar{c}$ , LOWER LINES  $b\bar{b}$   
 $m_c = 1.5 \text{ GeV}$     $m_b = 4.5 \text{ GeV}$

$$\int_0^1 dz g_1^{Q\bar{Q}}(z, Q^2) = 0$$

LO  
 NLO?



$$g_2^{Q\bar{Q}}(x, Q^2)$$



UPPER LINES:  $c\bar{c}$ , LOWER LINES  $b\bar{b}$   
 $m_c = 1.5 \text{ GeV}$      $m_b = 4.5 \text{ GeV}$

$$g_2^{\Gamma Q\bar{Q}}(x, Q^2) = -g_1^{\Gamma Q\bar{Q}}(x, Q^2) + \int_x^1 \frac{dz}{z} g_1^{\Gamma Q\bar{Q}}(z, Q^2)$$

## System : $g_1(x, Q^2), \partial g_1 / \partial t(x, Q^2)$

---

Leading Order :

$$K_{22}^{N(0)} = 0$$

$$K_{2d}^{N(0)} = -4$$

$$K_{d2}^{N(0)} = \frac{1}{4} \left( \gamma_{qq}^{N(0)} \gamma_{gg}^{N(0)} - \gamma_{qg}^{N(0)} \gamma_{gq}^{N(0)} \right)$$

$$K_{dd}^{N(0)} = \gamma_{qq}^{N(0)} + \gamma_{gg}^{N(0)}$$

Next-to-Leading Order : [W. Furmanski and R. Petronzio, Z. Phys. C 11 (1982) 293.]

$$K_{22}^{N(1)} = K_{2d}^{N(1)} = 0$$

$$K_{d2}^{N(1)} = \frac{1}{4} \left[ \gamma_{gg}^{N(0)} \gamma_{qq}^{N(1)} + \gamma_{gg}^{N(1)} \gamma_{qq}^{N(0)} - \gamma_{qg}^{N(1)} \gamma_{gq}^{N(0)} - \gamma_{qg}^{N(0)} \gamma_{gq}^{N(1)} \right]$$

$$- \frac{\beta_1}{2\beta_0} \left( \gamma_{qq}^{N(0)} \gamma_{gg}^{N(0)} - \gamma_{qg}^{N(0)} \gamma_{gq}^{N(0)} \right)$$

$$+ \frac{\beta_0}{2} C_{2,q}^{N(1)} \left( \gamma_{qq}^{N(0)} + \gamma_{gg}^{N(0)} - 2\beta_0 \right)$$

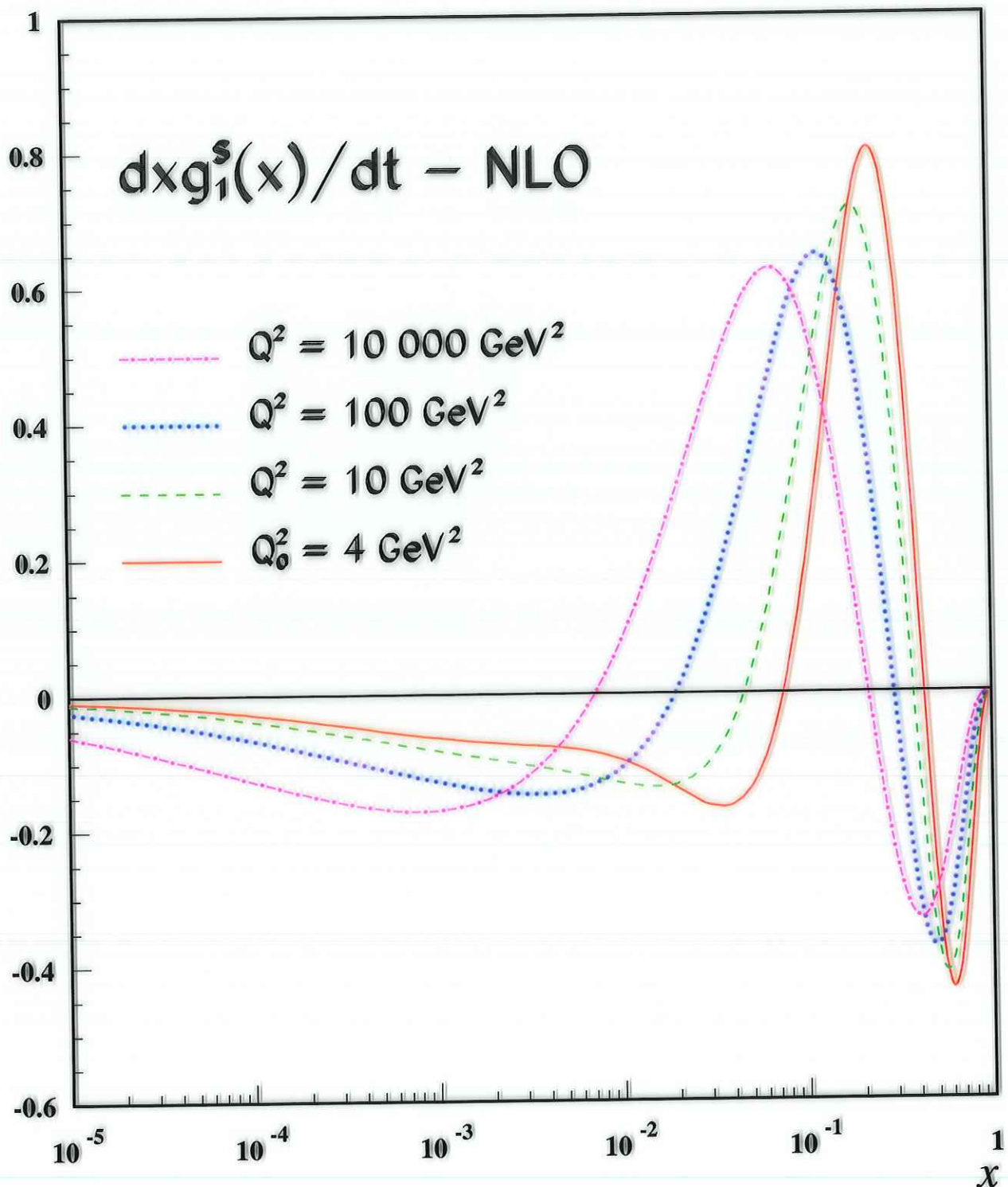
$$- \frac{\beta_0}{2} \frac{C_{2,g}^{N(1)}}{\gamma_{qg}^{N(0)}} \left[ \gamma_{qq}^{N(0)2} - \gamma_{qq}^{N(0)} \gamma_{gg}^{N(0)} + 2\gamma_{qg}^{N(0)} \gamma_{gq}^{N(0)} - 2\beta_0 \gamma_{qq}^{N(0)} \right]$$

$$- \frac{\beta_0}{2} \left( \gamma_{qq}^{N(1)} - \frac{\gamma_{qq}^{N(0)} \gamma_{qg}^{N(1)}}{\gamma_{qg}^{N(0)}} \right)$$

$$K_{dd}^{N(1)} = \gamma_{qq}^{N(1)} + \gamma_{gg}^{N(1)} - \frac{\beta_1}{\beta_0} \left( \gamma_{qq}^{N(0)} + \gamma_{gg}^{N(0)} \right) + 4\beta_0 C_{2,q}^{N(1)} - 2\beta_1$$

$$- \frac{2\beta_0}{\gamma_{qg}^{N(0)}} \left[ C_{2,g}^{N(1)} \left( \gamma_{qq}^{N(0)} - \gamma_{gg}^{N(0)} - 2\beta_0 \right) - \gamma_{qg}^{N(1)} \right]$$

## $\partial x g_1^S / \partial t(x, Q^2)$ and shift of $\Lambda_{QCD}^{(4)}$



Sc1 :  $\Lambda_{QCD}^{(4)} : 0.235 \rightarrow 0.223$  ,  $\alpha_s(M_Z^2) : 0.113 \rightarrow 0.112$

Sc2 :  $\Lambda_{QCD}^{(4)} : 0.240 \rightarrow 0.228$  ,  $\alpha_s(M_Z^2) : 0.114 \rightarrow 0.113$



## Fac. Scheme Invariant Combinations

---

- Instead of **PROCESS-INDEPENDENT SCHEME-DEPENDENT** Evolution Equations for **PARTONS** one may think of **PROCESS-DEPENDENT SCHEME-INDEPENDENT** Evolution Equations for **OBSERVABLES**,  $F_A, F_B$ .

- ⇒ The input densities are measured! Control over the input directly.
- ⇒ No  $\Delta G$ -Ansatz necessary.
- ⇒ A one parameter fit only –  $\Lambda_{QCD}$ .

Evolution Equations : [J. Blümlein, V. Ravindran, and W. L. van Neerven, Nucl. Phys **B586** (2000) 349.]

$$\frac{\partial}{\partial t} \begin{pmatrix} F_A^N \\ F_B^N \end{pmatrix} = -\frac{1}{4} \begin{pmatrix} K_{AA}^N & K_{AB}^N \\ K_{BA}^N & K_{BB}^N \end{pmatrix} \begin{pmatrix} F_A^N \\ F_B^N \end{pmatrix}$$

evolution variable :

$$t = -\frac{2}{\beta_0} \log \left( \frac{a_s(Q^2)}{a_s(Q_0^2)} \right)$$

- ⇒ The evolution kernels  $K_{IJ}^N$  are also Physical Quantities! The **Factorization Scheme Independence** holds order by order.

The **Renormalization Scale Dependence** disappears only with more higher orders.

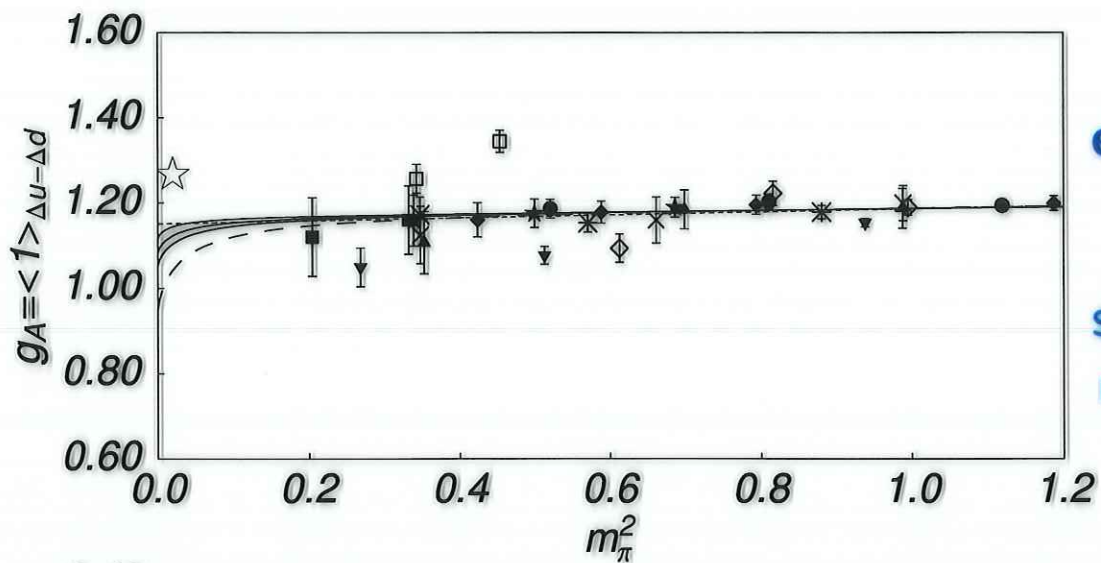
- ⇒ A possible choice:  $F_A = g_1$  and  $F_B = \partial g_1 / \partial t$ .

## 'Prediction' of Moments

---

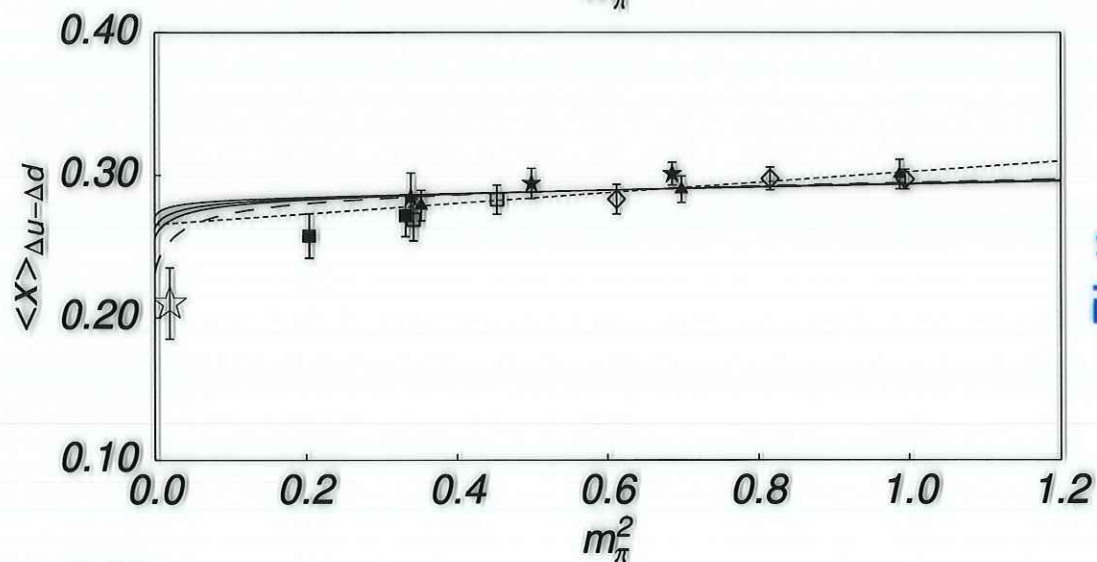
	$n$	QCD Scenario 1	
		value at $Q^2 = 4 \text{ GeV}^2$	value out of measured range
$\Delta u$	-1	$0.851 \pm 0.075$	$0.152   4\text{E}-4$
	0	$0.160 \pm 0.014$	$8\text{E}-4   3\text{E}-4$
	1	$0.055 \pm 0.006$	$1\text{E}-5   3\text{E}-4$
	2	$0.024 \pm 0.003$	$0   3\text{E}-4$
$\Delta d$	-1	$-0.415 \pm 0.124$	$-0.144   -7\text{E}-5$
	0	$-0.050 \pm 0.022$	$-7\text{E}-4   -6\text{E}-5$
	1	$-0.015 \pm 0.009$	$-1\text{E}-5   -5\text{E}-5$
	2	$-0.006 \pm 0.005$	$0   -5\text{E}-5$
$\Delta \bar{q}$	-1	$-0.074 \pm 0.017$	$-0.04   0$
	0	$-0.003 \pm 0.001$	$-2\text{E}-4   0$
	1	$-4\text{E}-4 \pm 1\text{E}-4$	$0   0$
	2	$-8\text{E}-5 \pm 2\text{E}-5$	$0   0$
$\Delta G$	-1	$1.026 \pm 0.549$	$0.04   1\text{E}-5$
	0	$0.184 \pm 0.103$	$5\text{E}-4   1\text{E}-5$
	1	$0.050 \pm 0.028$	$1\text{E}-5   1\text{E}-5$
	2	$0.017 \pm 0.010$	$0   1\text{E}-5$

# Lattice: The lowest moments of $\Delta u - \Delta d$

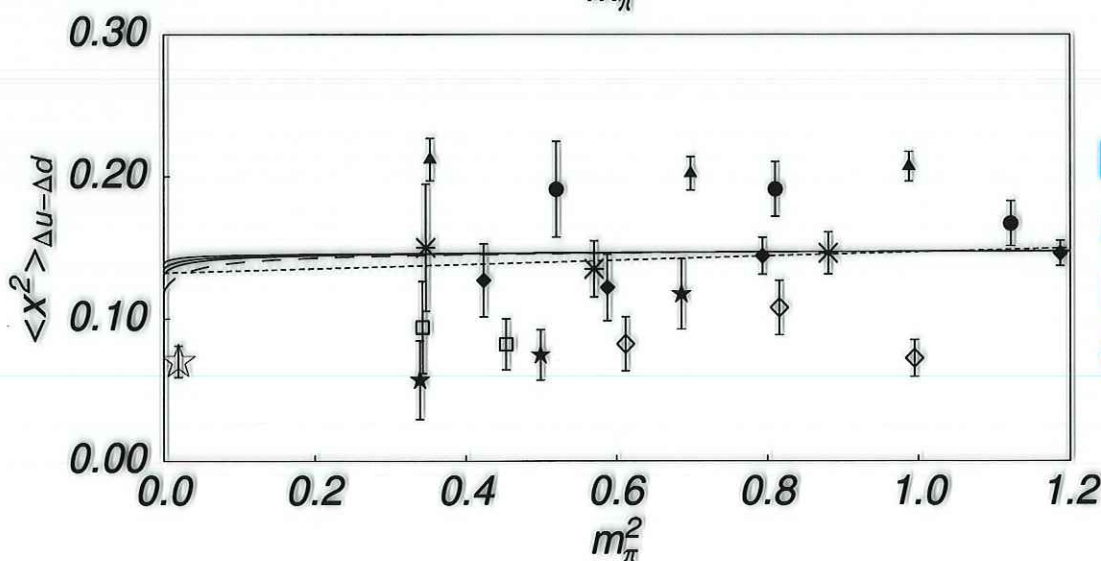


extrapolations:

short-dashed:  
naive linear



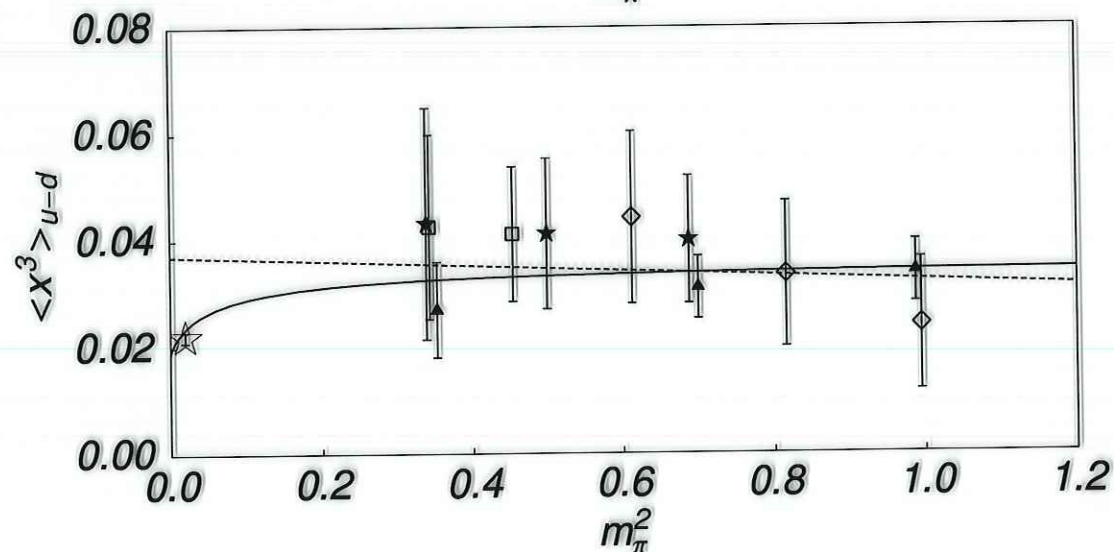
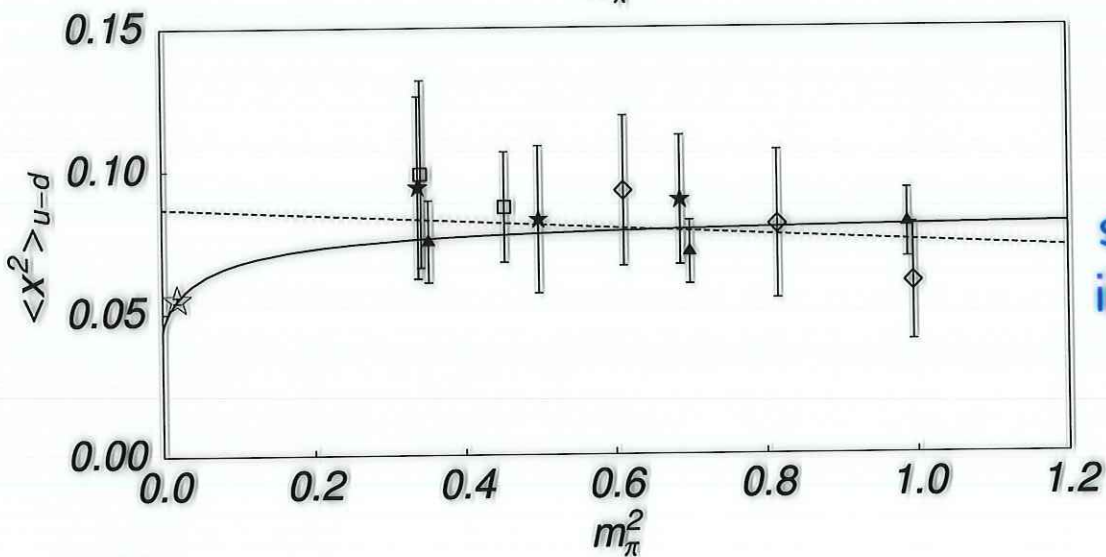
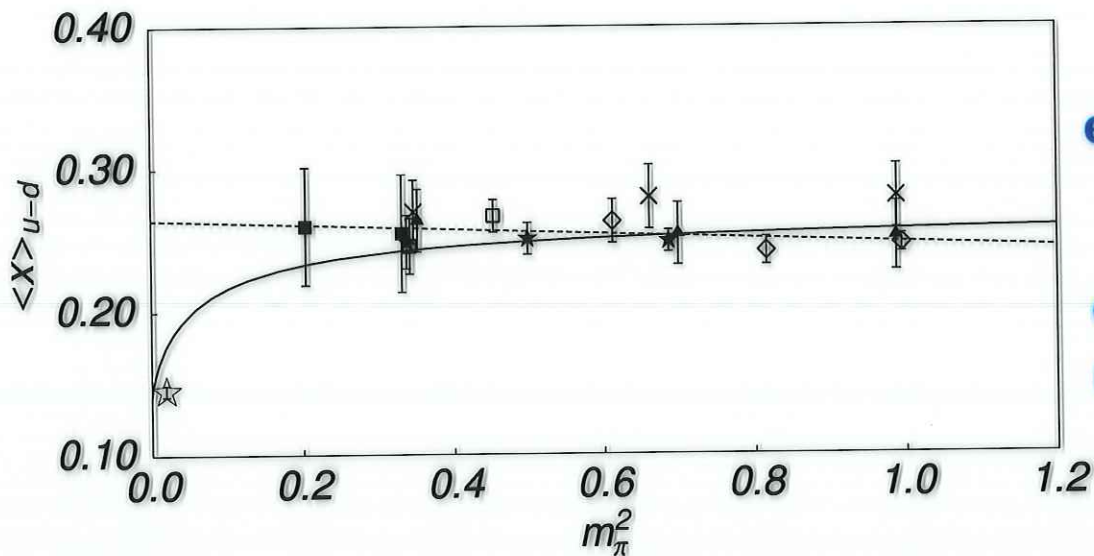
solid:  
improved chiral



long-dashed:  
no  $\Delta$  and  
LNA coeff.  
from  $\chi PT$



# Lattice: The lowest moments of $u - d$



Ref.: M. Detmold, W. Melnitchouk, A.W. Thomas; hep-lat/0206001.

## Comparison of Moments

$\Delta f$	$n$	QCD Scenario 1	lattice results	
		moment at $Q^2 = 4 \text{ GeV}^2$	QCDSF	LHPC/ SESAM
$\Delta u_v$	-1	$0.926 \pm 0.071$	0.889(29)	0.860(69)
	0	$0.163 \pm 0.014$	0.198(8)	0.242(22)
	1	$0.055 \pm 0.006$	0.041(9)	0.116(42)
$\Delta d_v$	-1	$-0.341 \pm 0.123$	-0.236(27)	-0.171(43)
	0	$-0.047 \pm 0.021$	-0.048(3)	-0.029(13)
	1	$-0.015 \pm 0.009$	-0.028(2)	0.001(25)
$\Delta u - \Delta d$	-1	$1.267 \pm 0.142$	1.14(3)	1.031(81)
	0	$0.210 \pm 0.025$	0.246(9)	0.271(25)
	1	$0.070 \pm 0.011$	0.069(9)	0.115(49)

$$\Rightarrow \Gamma_{\Delta f}(Q^2) = \int_0^1 x^{n+1} \Delta f(x, Q^2) dx$$

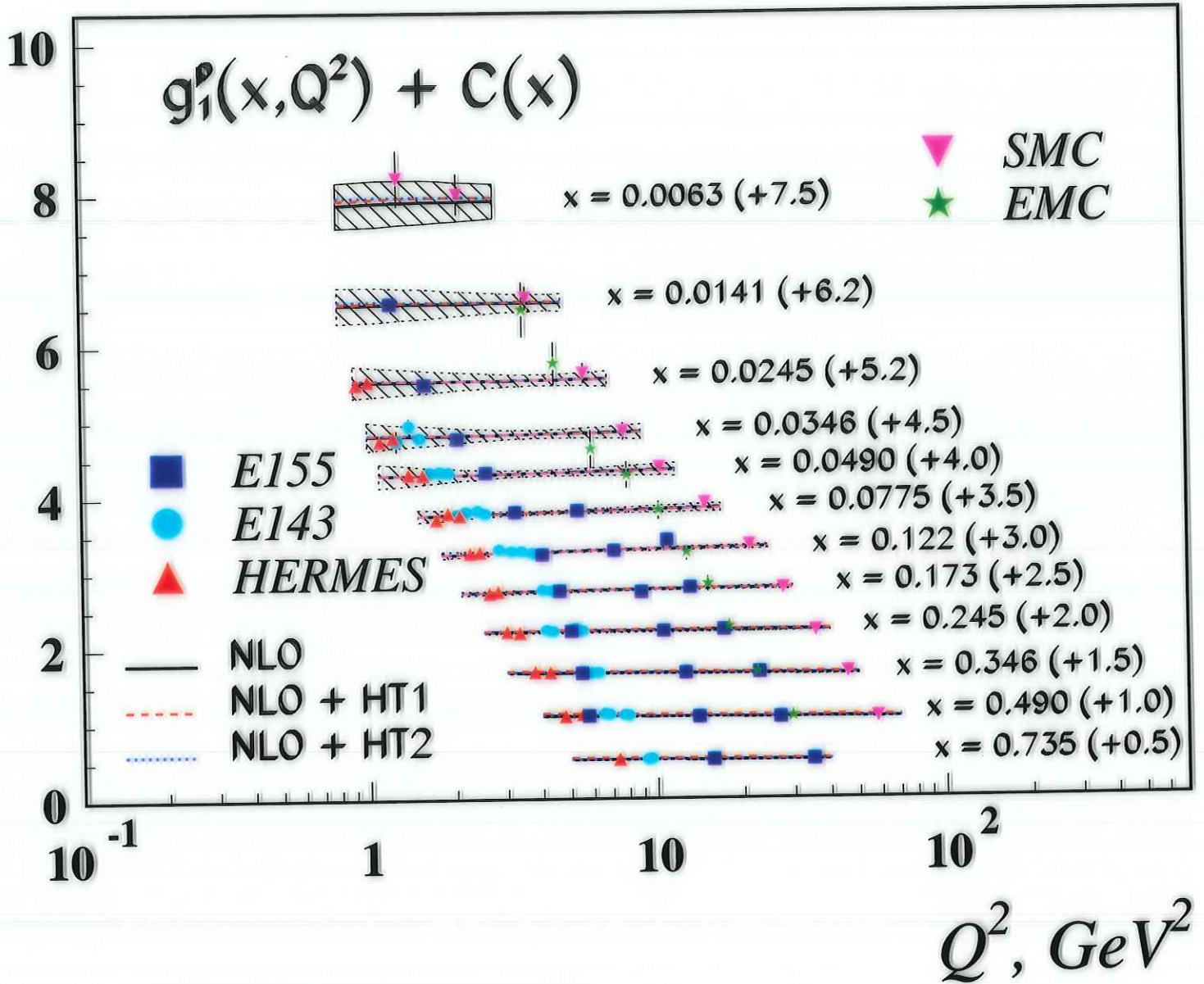
Lattice simulation: Scale  $\mu^2 = 1/a^2 \sim 4 \text{ GeV}^2$ . For the  $n = 0, 1$  values of the QCDSF Coll. no continuum extrapolation was performed.

[Refs: M.Göckeler et al., QCDSF Coll., Phys.Rev. **D53** (1996) 2317; Phys.Lett. **B414** (1997) 340; hep-ph/9711245; Phys.Rev. **D63** (2001) 074506; S.Capitani et al., Nucl.Phys.(Proc. Suppl.) **B79** (1999) 548; S.Güsken et al., SESAM Coll., hep-lat/9901009; D.Dolgov et al., LHPC and SESAM Coll., hep-lat/0201021.]

\* IMPROVED BY : SASAKI et al. hep-lat/0306007



# $g_1^p(x) + \text{Higher Twist} - \text{Scenario 1}$

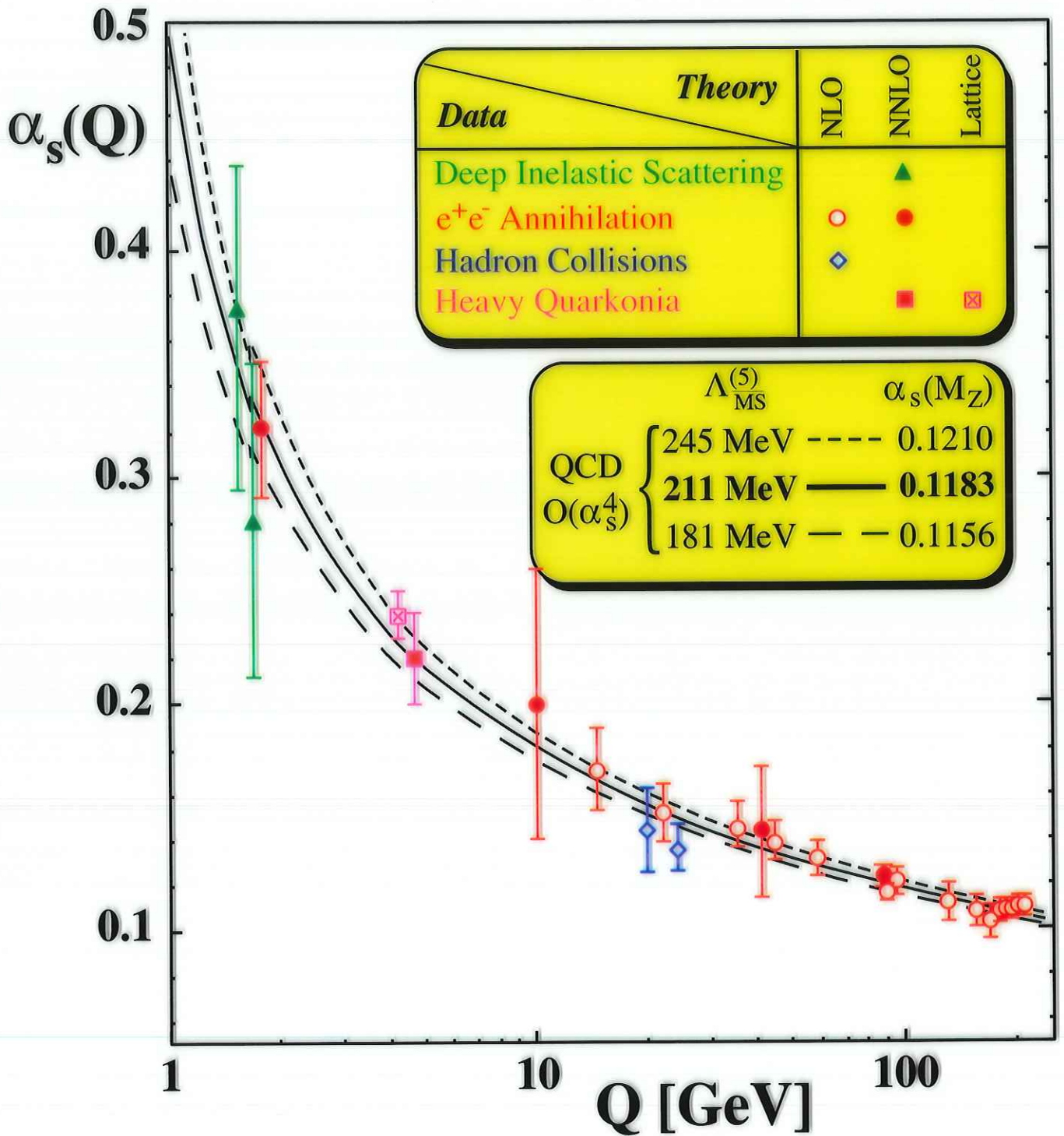


⇒ Hatched error band: Fully correlated  $1\sigma$  Gaussian error propagation through the evolution equation.

- Higher Twist Contribution:  $g_1(x, Q^2)[1 + HT(x, Q^2)]$ 
  - HT1:  $(1/Q^2)(x^a(1-x)^b)$
  - HT2:  $(1/Q^2)(a + bx + cx^2)$



# The QCD Running Coupling Constant



S. Bethke, 2002.

## 7+1 parameter NLO fit: $\Lambda_{QCD}^{(4)} \Rightarrow \alpha_s(M_Z^2)$

$\Lambda_{QCD}^{(4)}$ [Gev]	Scenario 1		Scenario 2	
	value	error	value	error
FS/RS=1.0/1.0	0.235	$\pm 0.053$	0.240	$\pm 0.060$
FS/RS=0.5/1.0	0.188	$- 0.047$	0.195	$- 0.045$
FS/RS=2.0/1.0	0.296	$+ 0.061$	0.298	$+ 0.058$
FS/RS=1.0/0.5	0.349	$+ 0.114$	0.363	$+ 0.123$
FS/RS=1.0/2.0	0.174	$- 0.061$	0.174	$- 0.066$

- Sc. 1:
 

$\alpha_s(M_Z^2) = 0.113$	$+0.004$	$+0.004$	$+0.008$
	$-0.004$	$-0.004$	$-0.005$
	(fit)	(fac)	(ren)
- Sc. 2:
 

$\alpha_s(M_Z^2) = 0.114$	$+0.004$	$+0.004$	$+0.008$
	$-0.005$	$-0.004$	$-0.006$

- SMC:  $0.121 \pm 0.002(stat) \pm 0.006(syst + theor)$

E154:  $0.108 - 0.116$  (bad for  $\geq 0.120$ )

ABFR:  $0.120$   $+0.004$  (exp)  $+0.009$  (theor)  
 $-0.005$   $-0.006$

$\Rightarrow$  world average (PDG):  $0.118 \pm 0.002$

$\Rightarrow$  H1 + BCDMS data:  $0.1150 \pm 0.0017(exp)$   $+0.0009$  (model)  $\pm 0.0050(thy)$   
 $-0.0005$  [Eur.Phys.J. C21(2001)33]

## ANGULAR MOMENTUM

$$\frac{1}{2} = \frac{1}{2} \Delta \Sigma + L_q + J_g$$

$$J_g = \Delta q + "L_g"$$

DECOMPOS. NOT GAUGE INV.

$$\frac{d}{d \ln p^2} \begin{pmatrix} J_q \\ J_g \end{pmatrix} = \frac{\alpha_s}{2\pi} \frac{1}{9} \begin{pmatrix} -16 & 3N_f \\ 16 & -3N_f \end{pmatrix} \begin{pmatrix} J_q \\ J_g \end{pmatrix}$$

$$p^2 \rightarrow \infty \quad J_q \rightarrow \frac{1}{2} \frac{3N_f}{16 + 3N_f} \simeq 0.214 (N_f = 4)$$

$$J_g \rightarrow \frac{1}{2} \frac{16}{16 + 3N_f} \simeq 0.286 (N_f = 4)$$

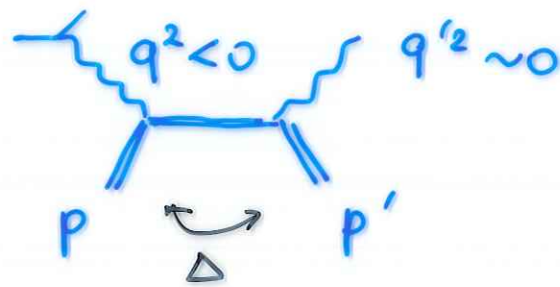
MEASUREMENT:

NEED 2<sup>nd</sup> VECTOR

→ NON-FORWARD SCATTERING



DVCS



$$\int \frac{d\lambda}{2\pi} e^{i\lambda x} \langle p' | \bar{\psi}(-\lambda n/2) \gamma^\mu \psi(\lambda n/2) | p \rangle$$

$$= H(x, \Delta^2, \xi) \bar{u}(p') \gamma^\mu u(p)$$

$$+ E(x, \Delta^2, \xi) \bar{u}(p') \frac{i\sigma^{\mu\nu} \Delta_\nu}{2M} u(p) + \dots$$

$$J_q = \frac{1}{2} \int_{-1}^{+1} dx \times [H_q(x, \Delta^2=0, \xi) + E_q(x, \Delta^2=0, \xi)]$$

↑  
extrapolation.

$\xi$  - non forwardness.

## Conclusions

---

- AN LO AND NLO QCD ANALYSIS OF THE CURRENT WORLD-DATA OF POLARIZED STRUCTURE FUNCTIONS WAS PERFORMED.
- **NEW PARAMETRIZATIONS** OF THE **PARTON DENSITIES** INCLUDING THEIR **FULLY CORRELATED  $1\sigma$  ERROR BANDS** WERE DERIVED. THEY ARE AVAILABLE VIA A FAST FORTRAN PROGRAM FOR THE RANGE:

$$1 \text{ GeV}^2 < Q^2 < 10^6 \text{ GeV}^2 \text{ AND } 10^{-9} < x < 1.$$

- THE FOLLOWING VALUES FOR  $\alpha_s(M_Z^2)$  WERE OBTAINED:

- SCENARIO 1:

$$\alpha_s(M_Z^2) = 0.113 \begin{array}{l} +0.004 \\ -0.004 \end{array} \text{ (fit)} \begin{array}{l} +0.004 \\ -0.004 \end{array} \text{ (fac)} \begin{array}{l} +0.008 \\ -0.005 \end{array} \text{ (ren)},$$

- SCENARIO 2:

$$\alpha_s(M_Z^2) = 0.114 \begin{array}{l} +0.004 \\ -0.005 \end{array} \text{ (fit)} \begin{array}{l} +0.004 \\ -0.004 \end{array} \text{ (fac)} \begin{array}{l} +0.008 \\ -0.006 \end{array} \text{ (ren)},$$

COMPATIBLE WITH RESULTS FROM OTHER QCD ANALYSES AND WITH THE WORLD AVERAGE.

## Conclusions (cont'd)

---

- FIRST STEPS IN A FACTOR. SCHEME INVARIANT QCD EVOLUTION BASED ON THE STRUCTURE FUNCTION  $g_1(x, Q^2)$  AND  $\partial g_1(x, Q^2)/\partial \log Q^2$  WERE PERFORMED YIELDING SIMILAR RESULTS FOR  $\alpha_s(M_Z^2)$ .

SUCH AN ANALYSIS IS A VERY PROMISING WAY TO PROCEED IN THE FUTURE, SINCE IT ALLOWS TO EXTRACT  $\Lambda_{\text{QCD}}$  FIXING ALL THE INPUT DISTRIBUTIONS BY DIRECT MEASUREMENT.

- COMPARING THE QCD LOW MOMENTS WITH VALUES FROM LATTICE SIMULATIONS THE ERRORS IMPROVED DURING RECENT YEARS AND THE VALUES BECAME CLOSER. HOWEVER, MORE WORK HAS YET TO BE DONE IN THE FUTURE ON SYSTEMATIC EFFECTS AND EVEN MORE PRECISE EXPERIMENTAL DATA ARE WELCOME TO IMPROVE PRECISION.



## 9. Future Avenues

### HERA:

- Collect high luminosity for  $F_2(x, Q^2)$ ,  $F_2^{cc}(x, Q^2)$ ,  $g_2^{cc}(x, Q^2)$ , and measure  $h_1(x, Q^2)$ .
- Measure :  $F_L(x, Q^2)$ . This is a key-question for HERA.

### RHIC & LHC:

- Improve constraints on gluon and sea-quarks: polarized and unpolarized.

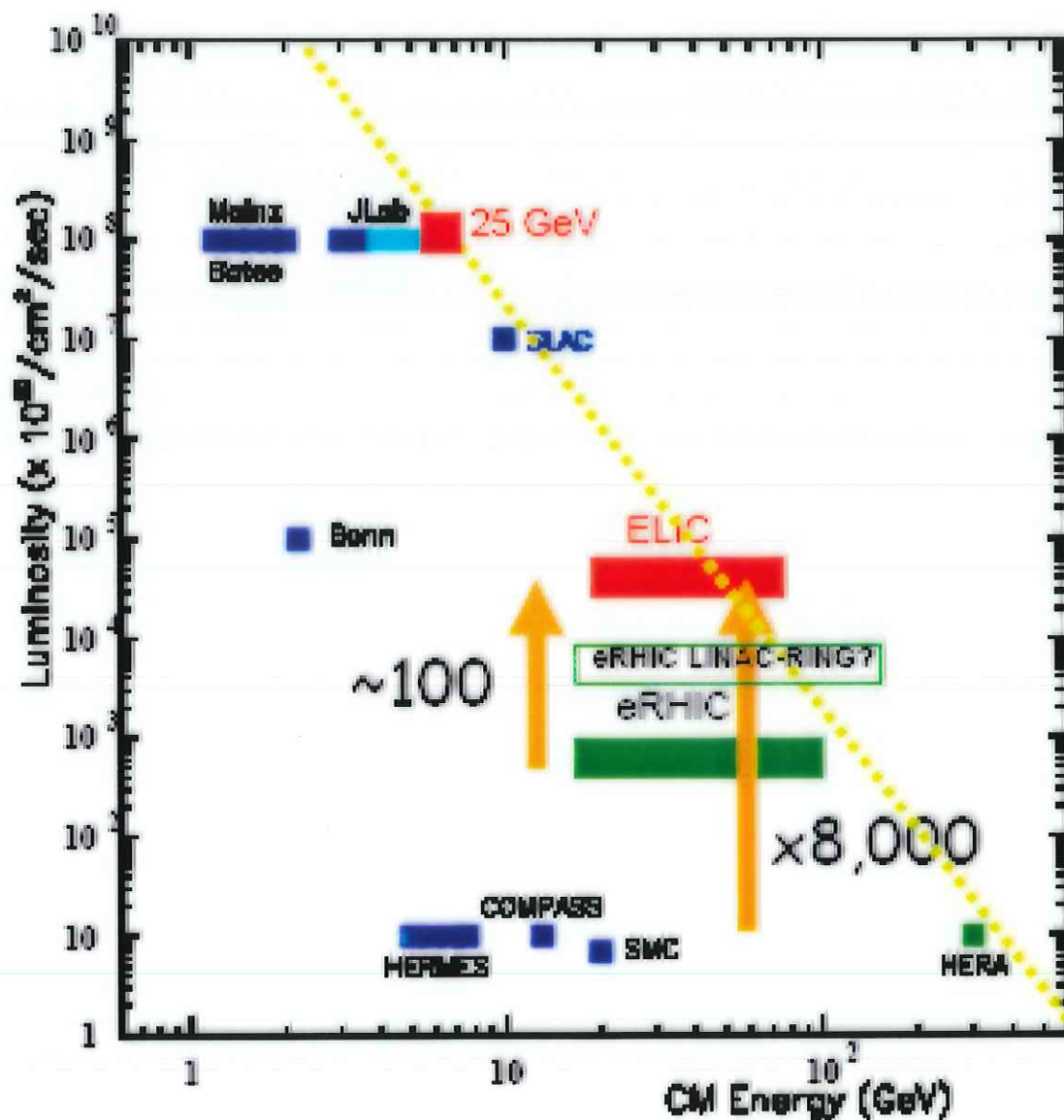
### JLAB:

- High precision measurements in the large  $x$  domain at unpolarized and polarized targets; supplements HERA's high precision measurements at small  $x$ .

## ELIC:

- High precision measurements in the medium  $x$  domain; both unpolarized and polarized

## THE QUEST FOR LARGE LUMINOSITY !



- What is the correct value of  $\alpha_s(M_z^2)$ ?  $\overline{\text{MS}}$ -analysis vs. scheme-invariant evolution helps. Compare non-singlet and singlet analysis; careful treatment of heavy flavor. [Theory & Experiment]
- Flavor Structure of Sea-Quarks: More studies needed. [All Experiments]
- Revisit polarized data upon arrival of the 3-loop anomalous dimensions; NLO heavy flavor contributions needed. [Theory]
- QCD at Twist 3:  $g_2(x, Q^2)$ , semi-exclusive Reactions [High Precision polarized experiments, JLAB, EIC]
- Comparison with Lattice Results:  $\alpha_s$ , Moments of Parton Distributions, Transversity, Angular Momentum.
- Calculation of more hard scattering reactions at the 3-loop level: ILC, LHC
- Further perfection of the mathematical tools:  
 $\implies$  Algorithmic simplification of Perturbation theory in higher orders.
- Even higher order corrections needed ?
- DIS AS A TERRITORY FOR PERTURBATIVE AND NON-PERTURBATIVE PRECISION CALCULATIONS  
 $\implies$  B3